

***r*-process enrichment of the lightest galaxies by fast merging neutron star binaries**

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The observed statistics of r -process enrichment in the local universe points to the need for fast merging neutron star mergers!

How to study the sources of r -process elements in the early universe?

observe old (metal poor) stars that are
r-process enhanced


Where are r -process metal poor stars?



MW's halo



Ultra Faint
Dwarf
galaxies



Ultra Faint
Dwarf
galaxies

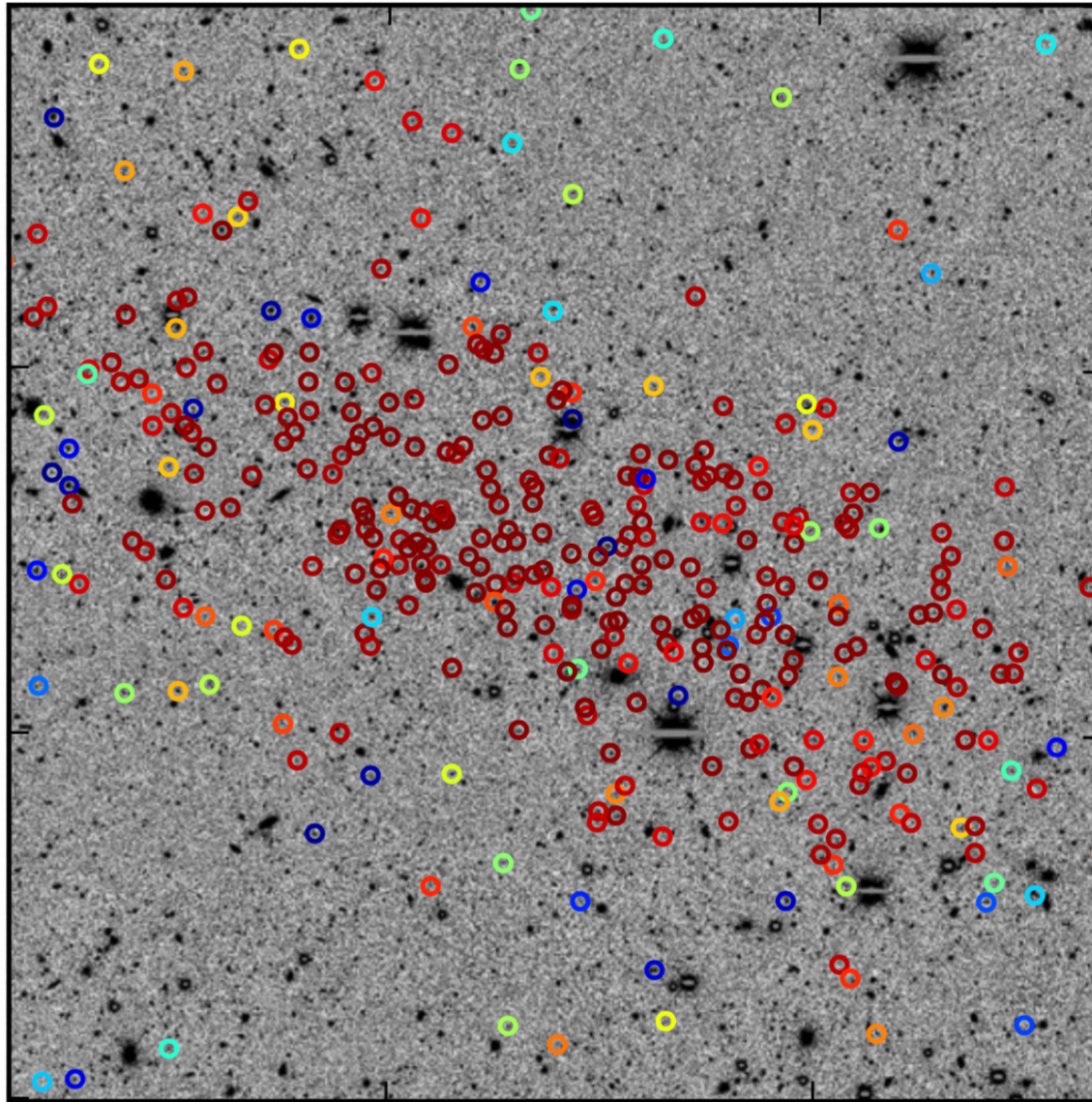
Reticulum II



Bechtol et al. 2015

UFDs: Dead, dark matter dominated,
least luminous, least chemically evolved systems

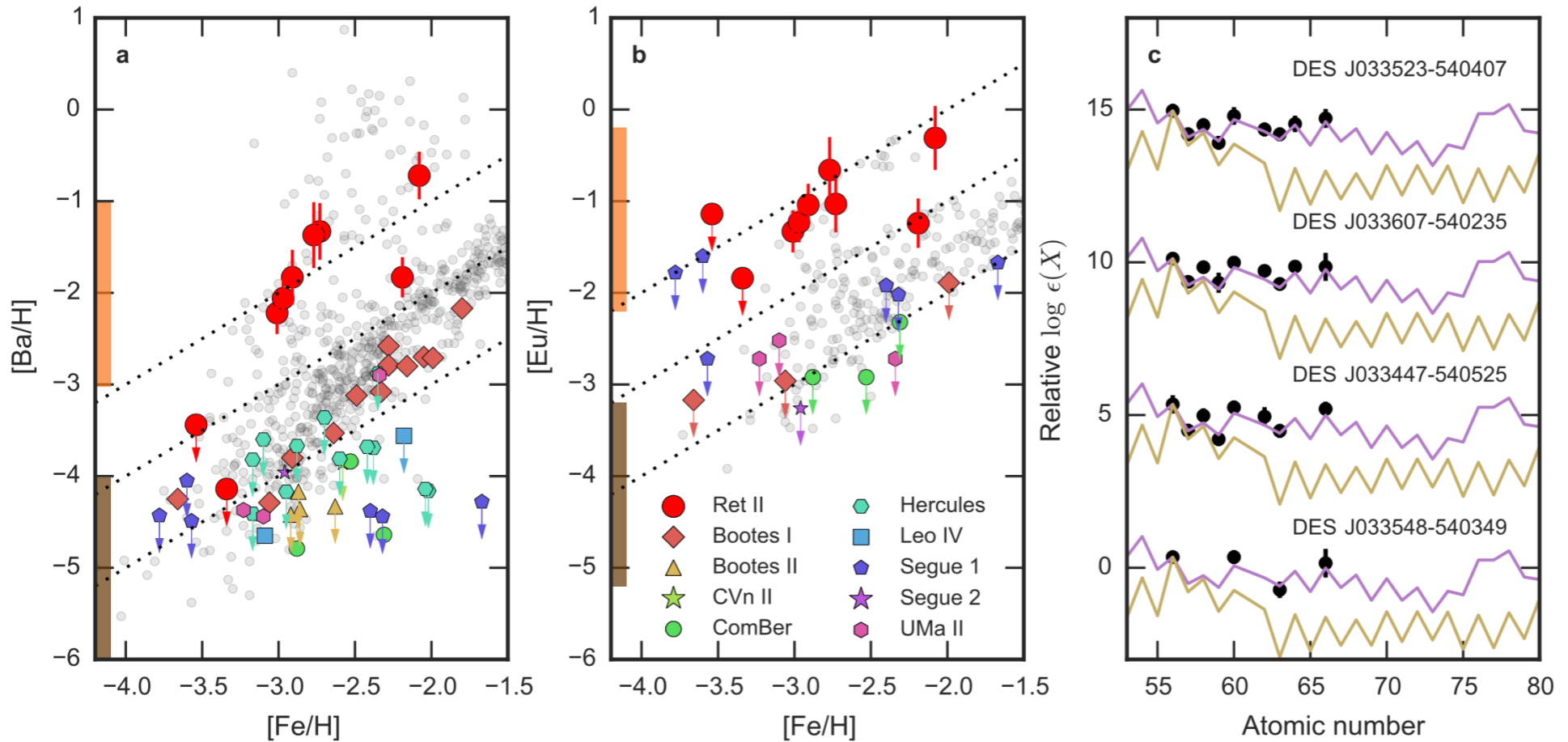
Reticulum II



Bechtol et al. 2015

UFDs: Dead, dark matter dominated,
least luminous, least chemically evolved systems

Reticulum II r -process abundance

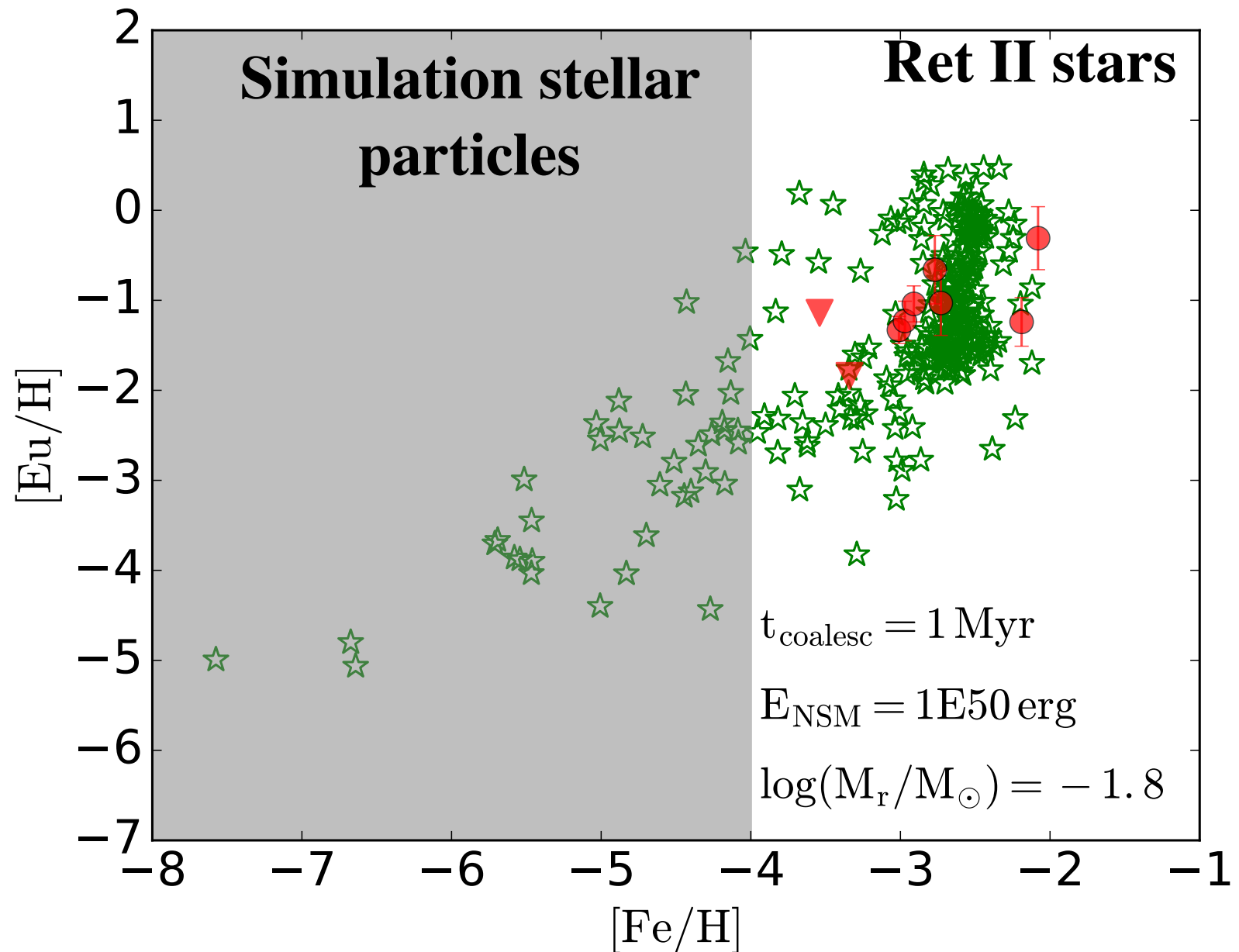


Ji et al. (2016)

Zoom cosmological simulation of a MW type
halo, tracking r -process enrichment

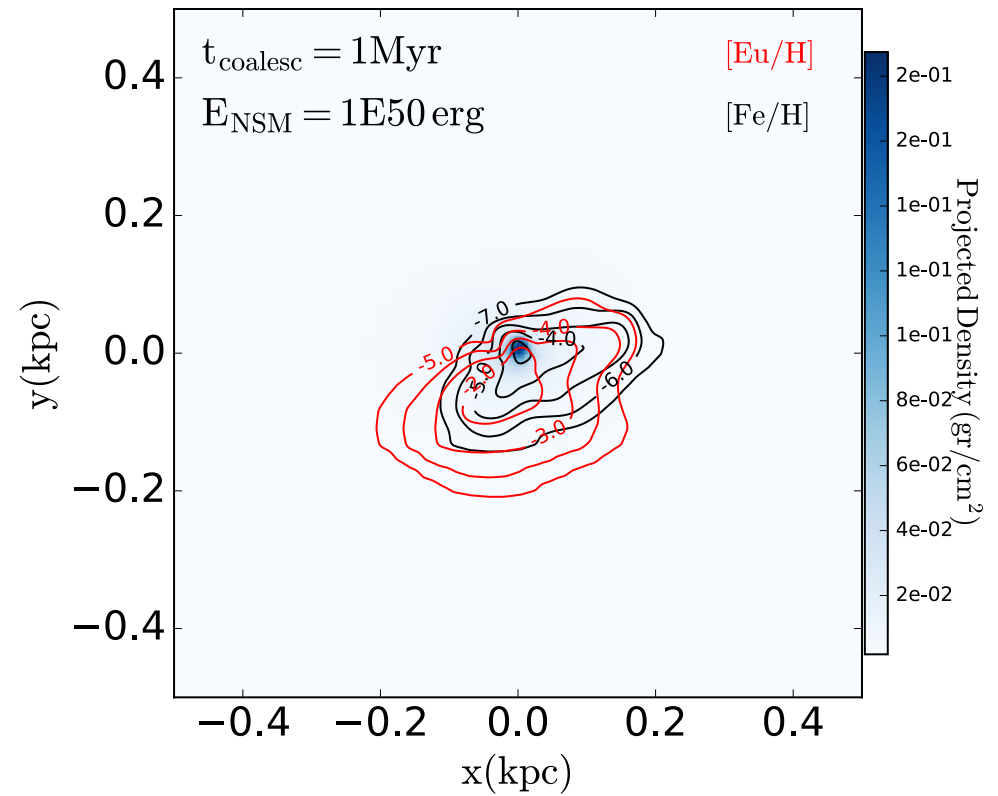
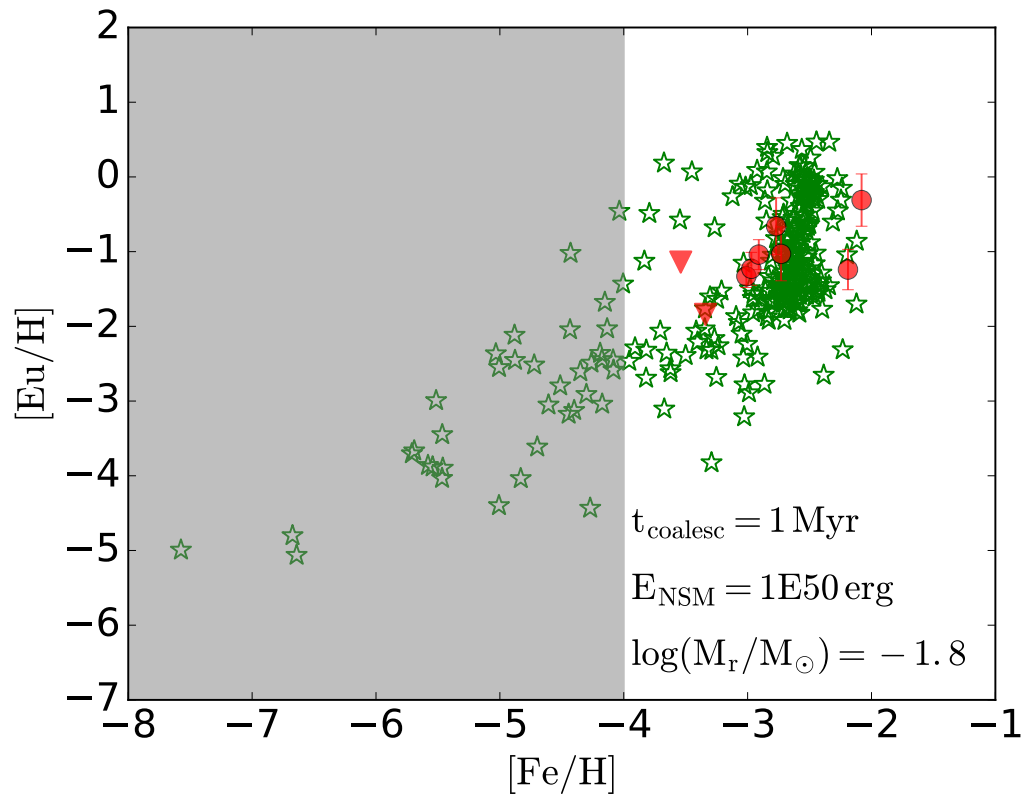
Results from zoom cosmological simulation of an ultra faint galaxy

Comparing the results to Ret II



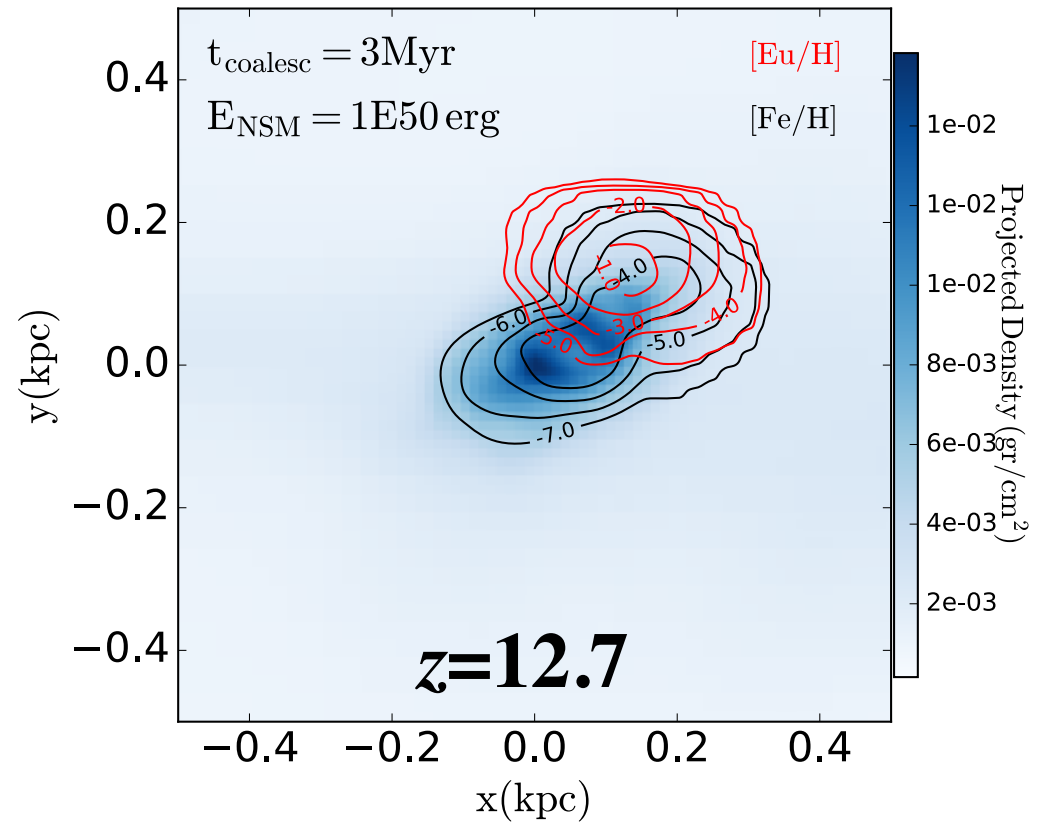
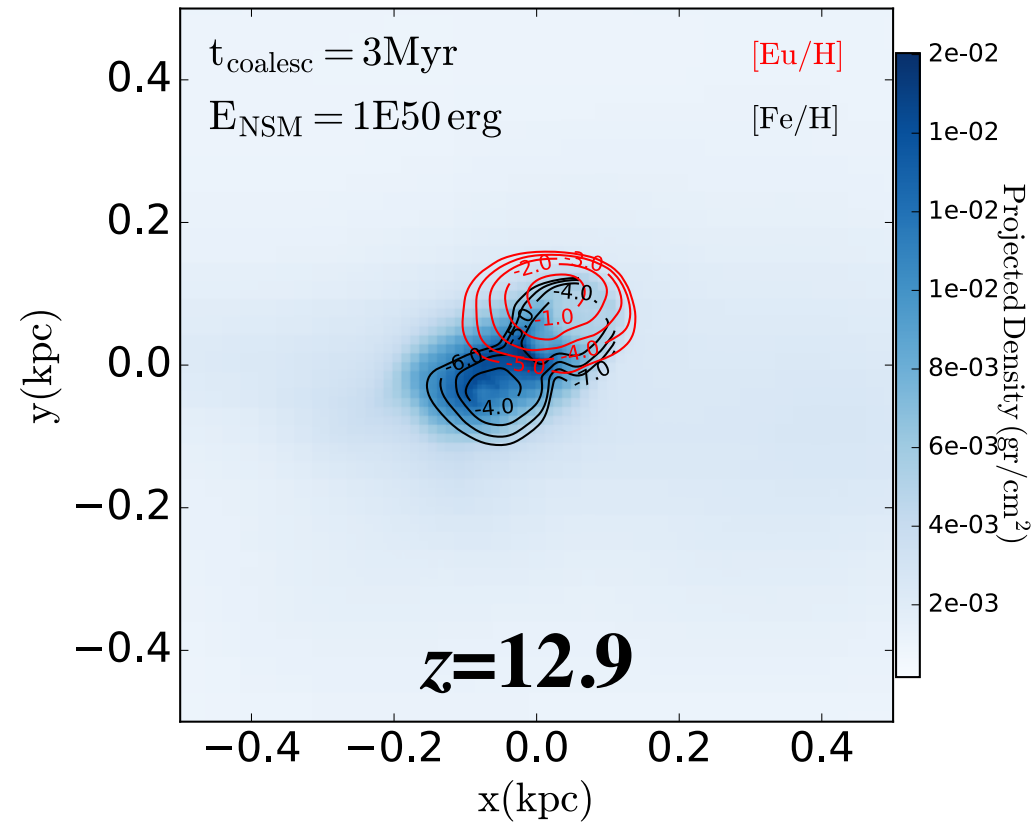
MTS & Scannapieco (2017)

Comparing the results to Ret II

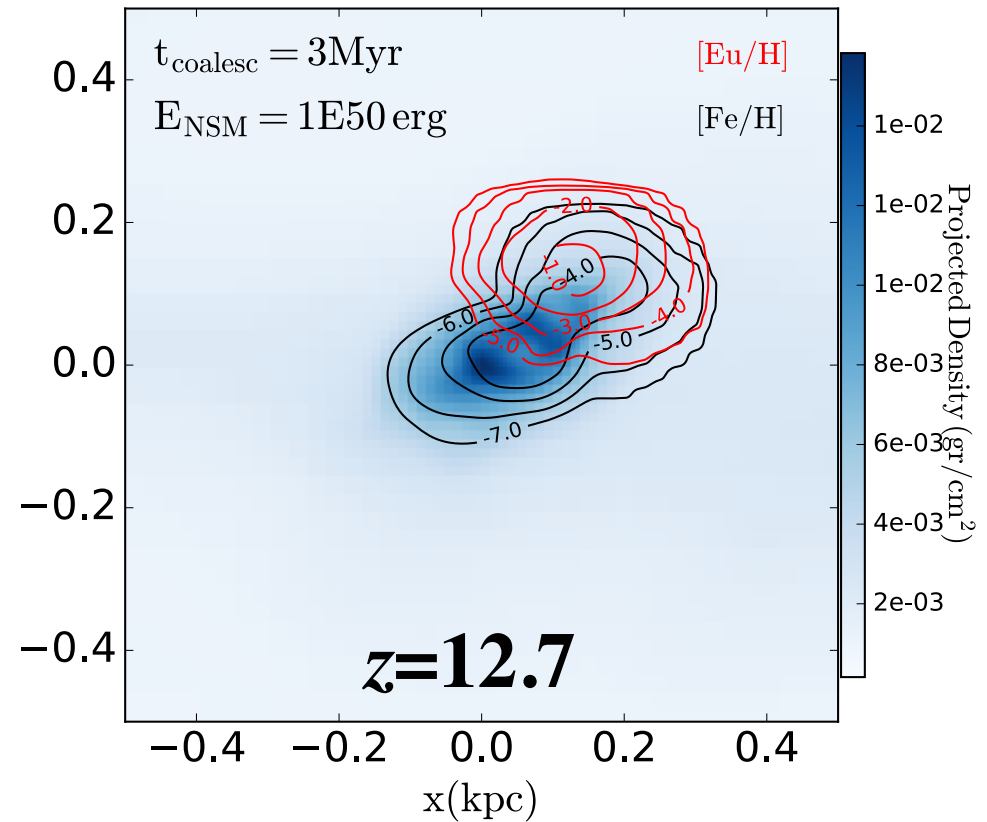
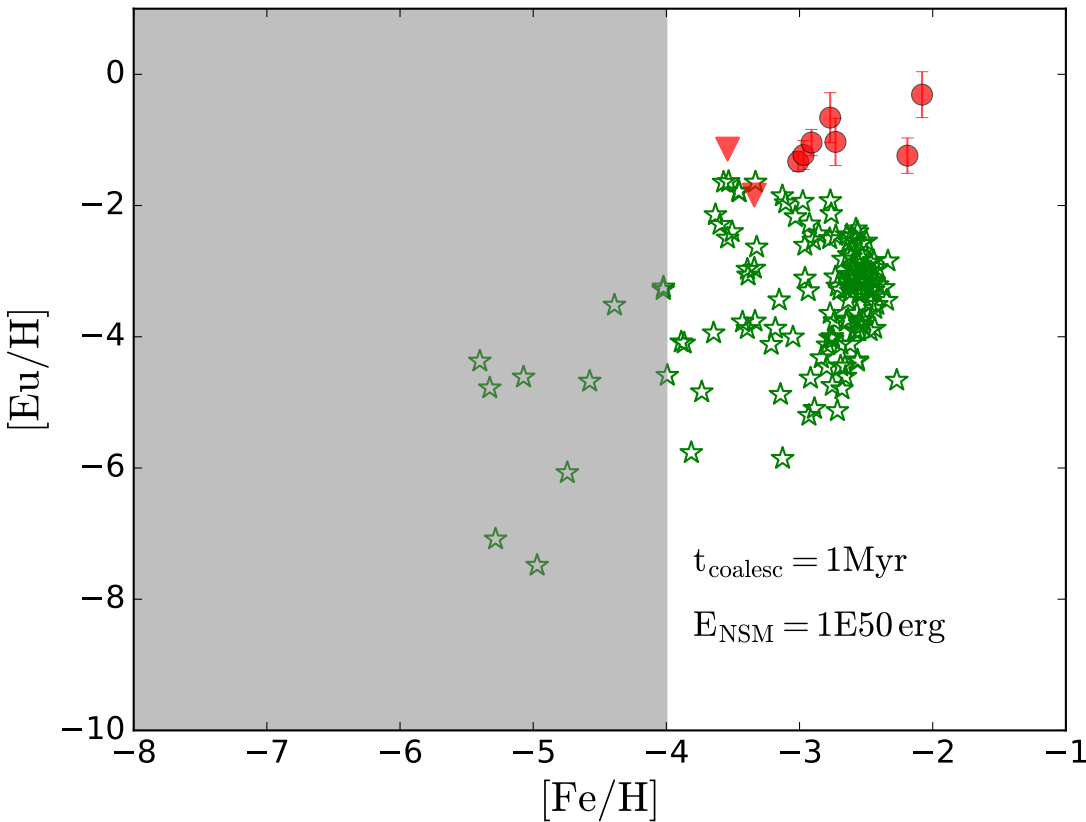


MTS & Scannapieco (2017)

Off-center explosion

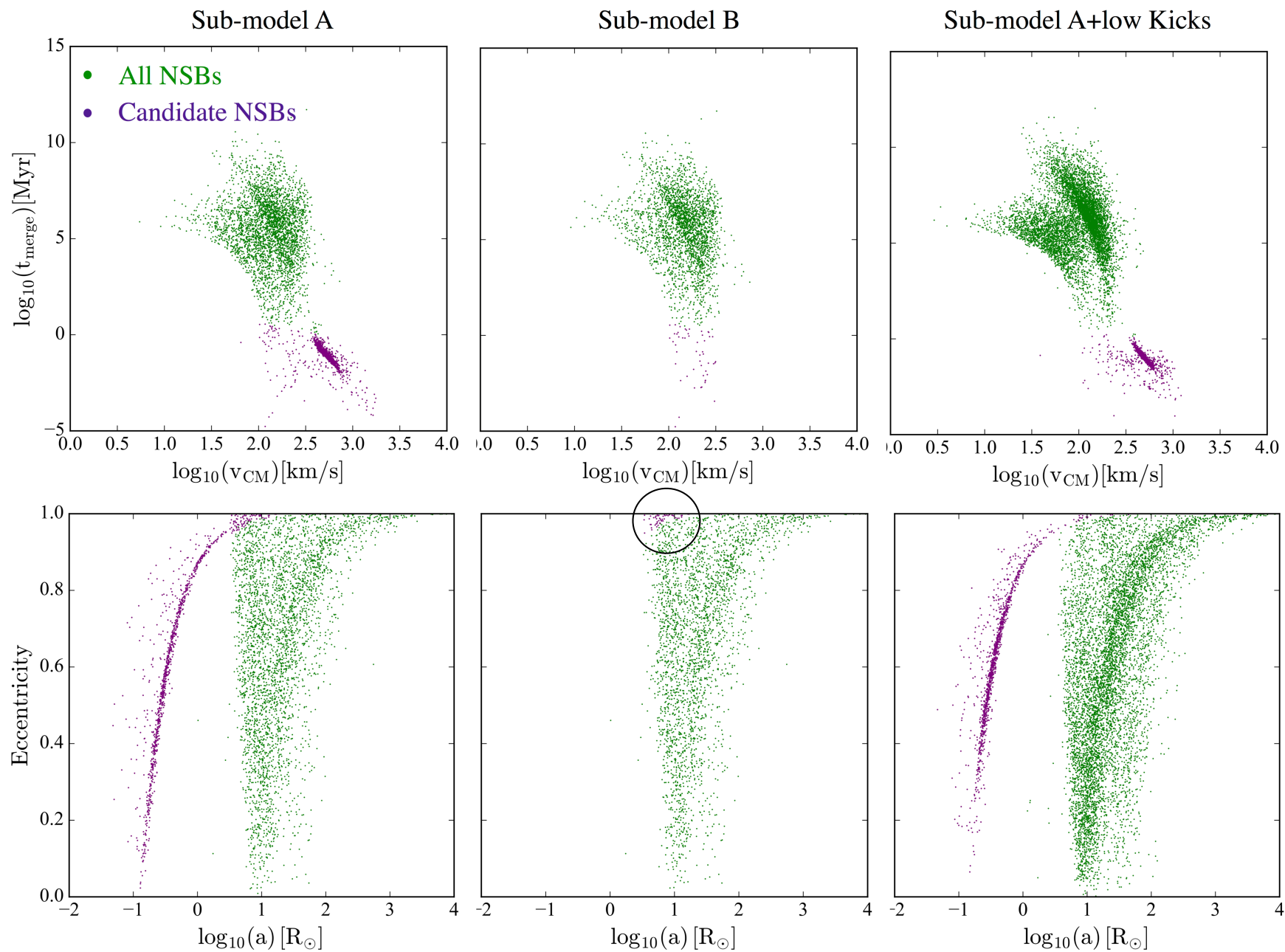


Off-center explosion




How can a NSM event enrich a galaxy?

Results from a distance cut on the NS
binaries.



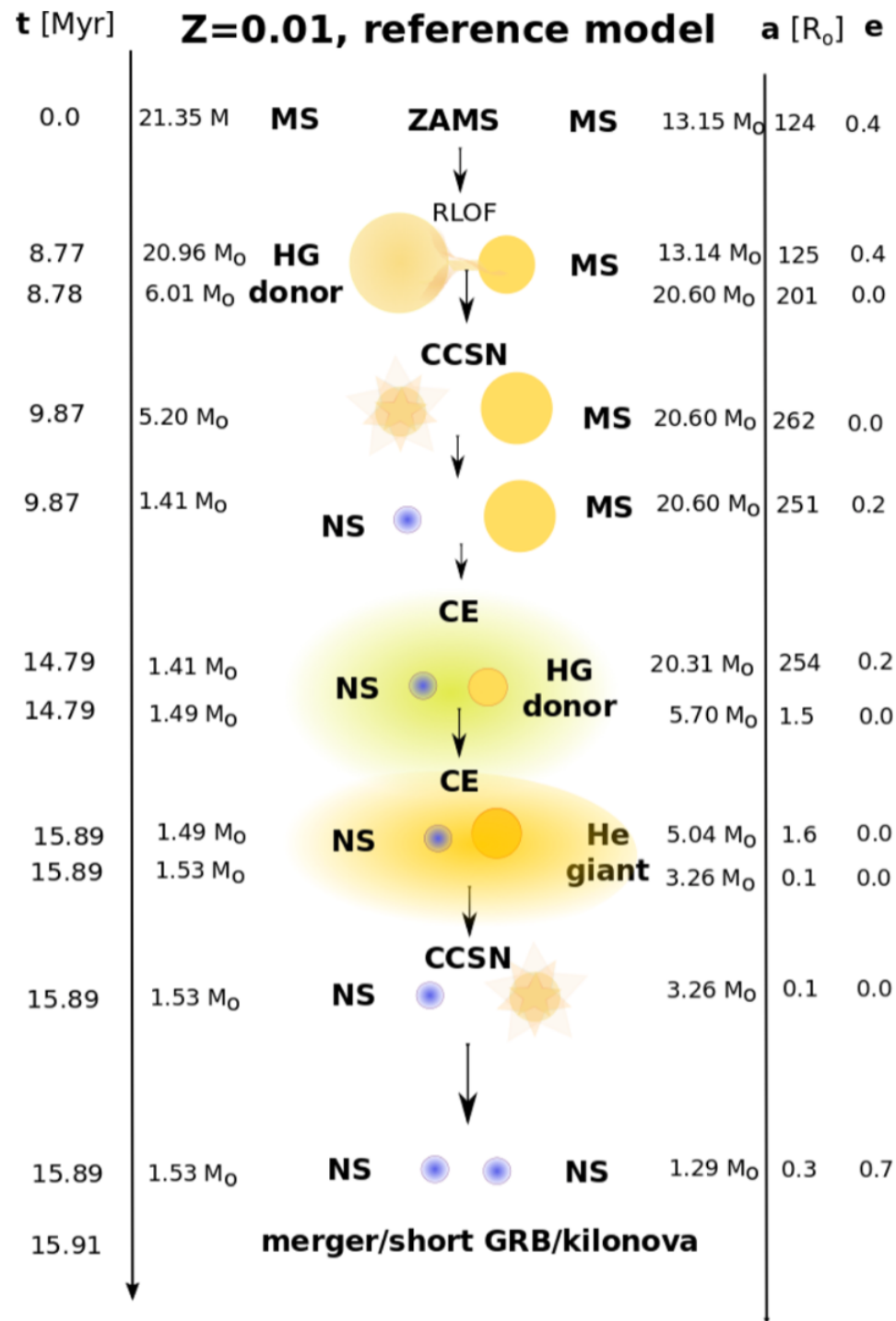
MTS, Ramirez-Ruiz et al. (2018)

Fast merging channel of NS binaries

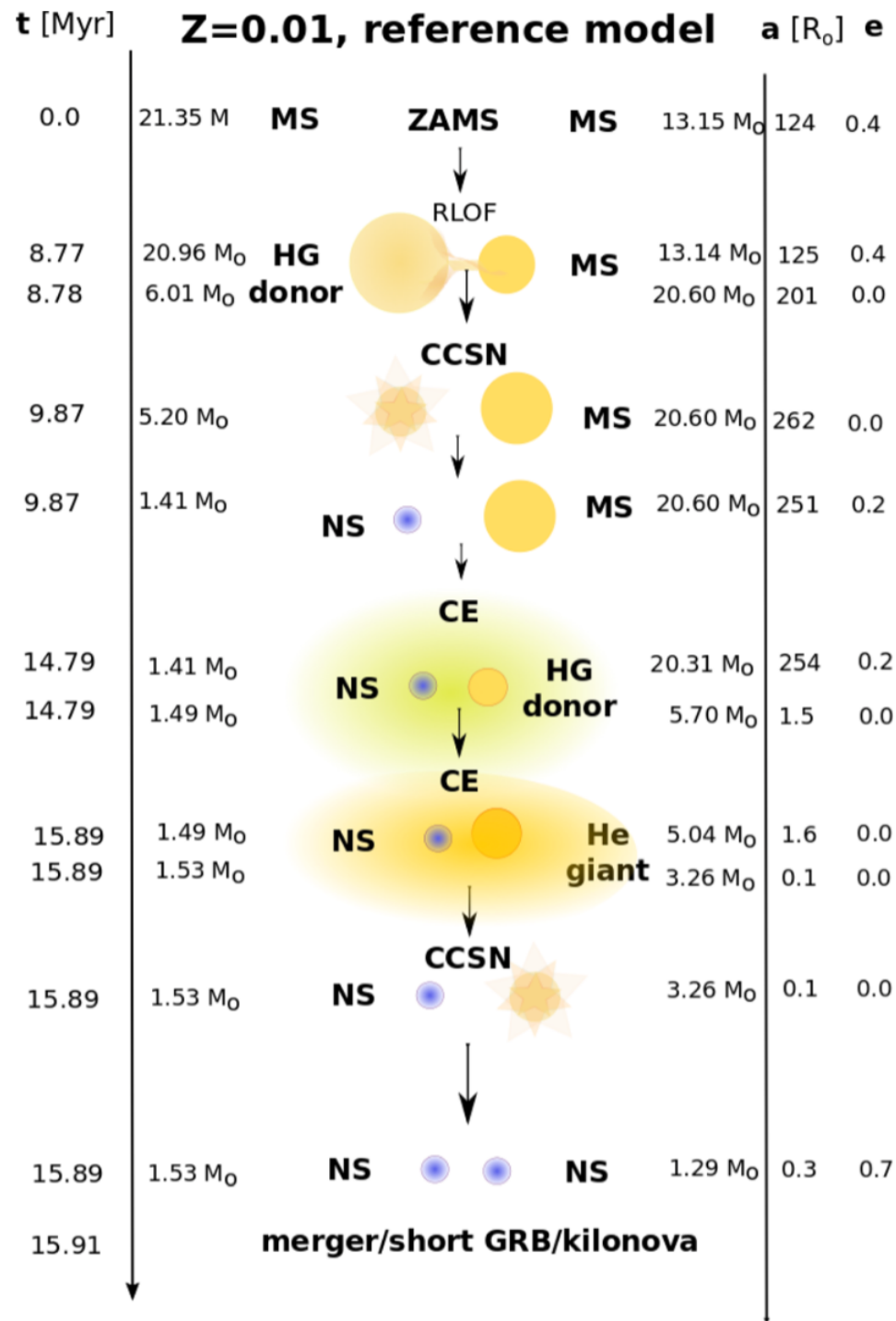


Highly
eccentric
orbits

Case BB
unstable
mass
transfer



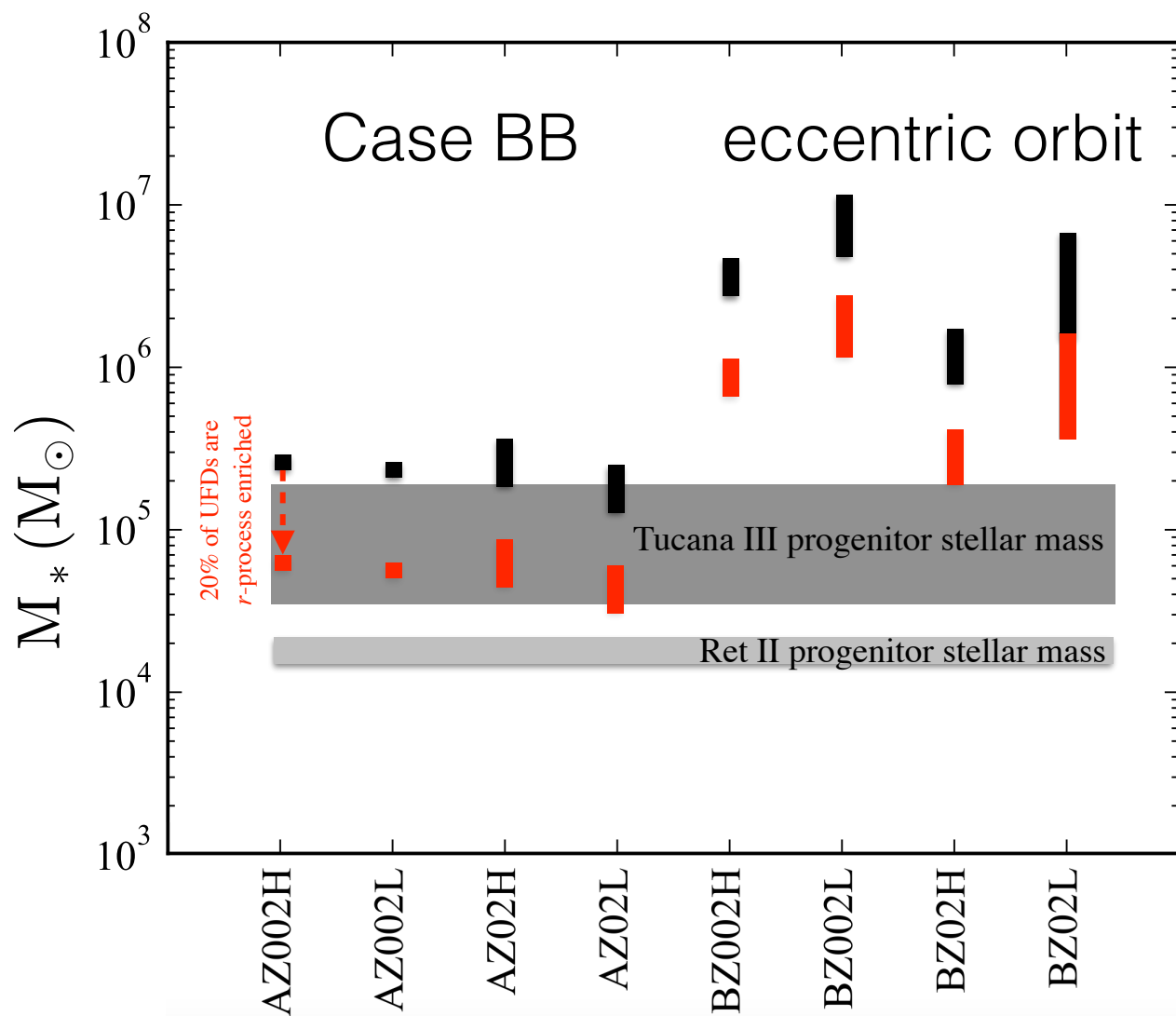
Case BB
unstable
mass
transfer



Case BB
unstable
mass
transfer

No CE simulation
successfully
ejects the envelope!

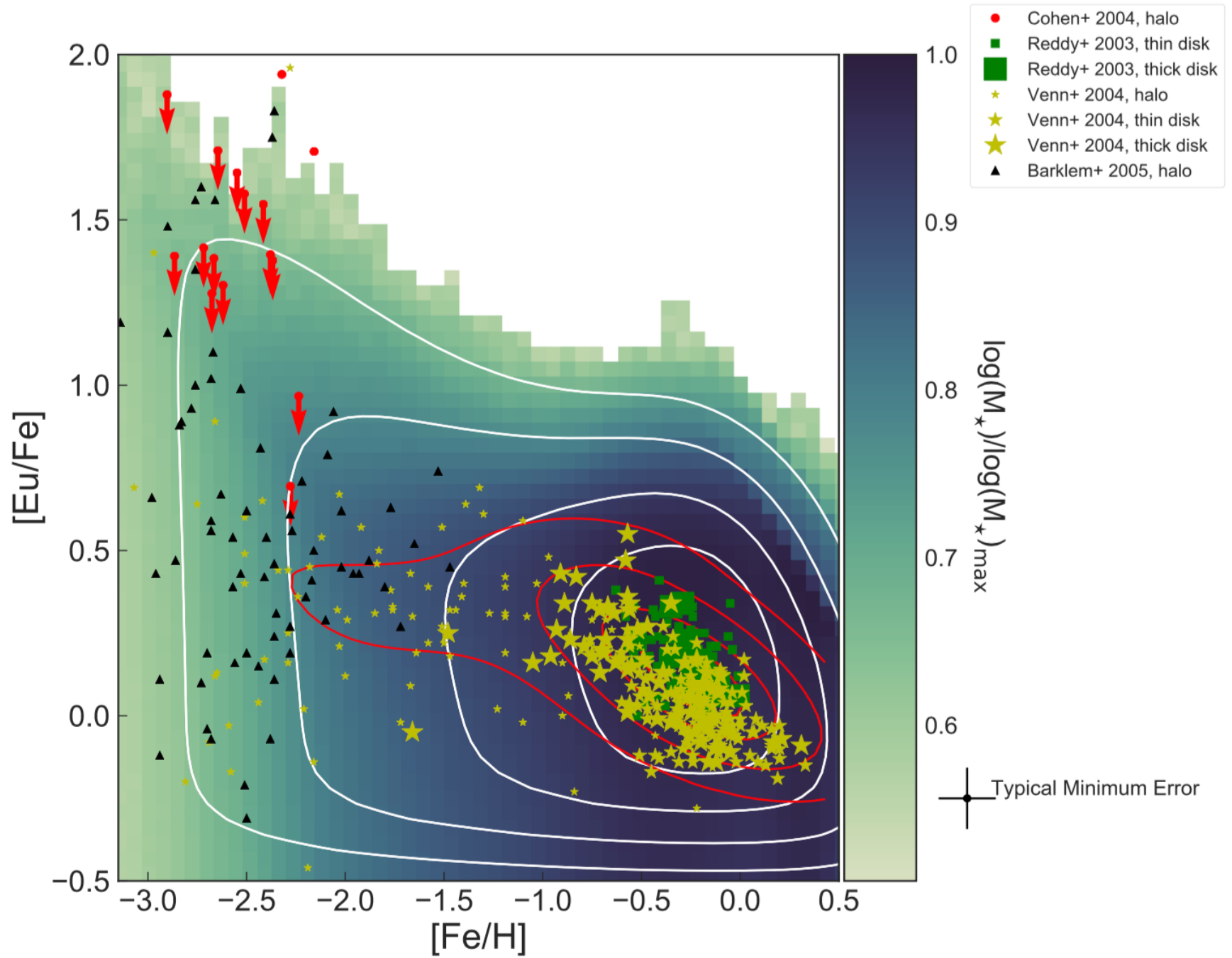
Chruslinska+18



MTS, Ramirez-Ruiz et al. (to be submitted)

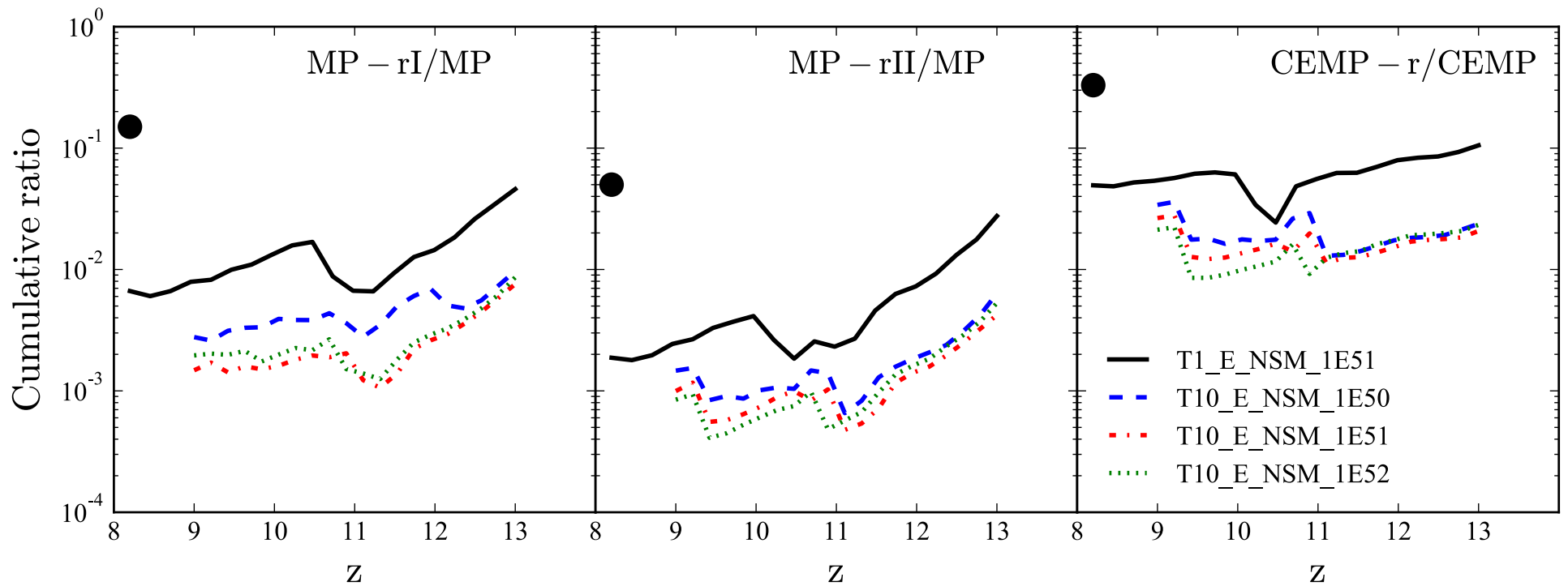
What is next?

Perform the same simulations on a suite of MW type halos to obtain a robust halo-to-halo scatter



Naiman et al., 2018

Performing cosmological zoom simulation of r -process enrichment on a MW type halo



MTS, sarmento & Scannapieco., in prep

Summary

A single NSM event in star formation history of a UFD is compatible with Ret II observations.

***r*-process enrichment efficiency is highly dependent on the location of NSM event, therefore natal kicks play a crucial role.**

Fast merging channels are crucial to explain Ret II like systems.

Case BB unstable mass transfer

Highly eccentric orbits