JINA: Center for the Evolution of the Elements

Neutrino Cooling in the Accreted Neutron Star Ocean

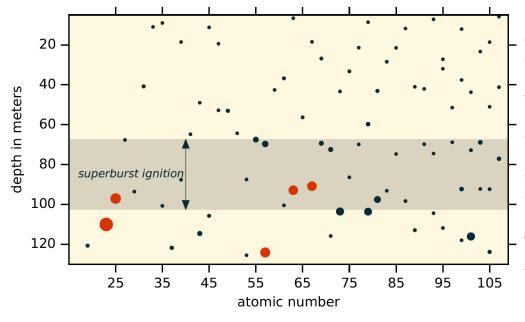


Many neutron stars accrete hydrogen- and helium-rich matter from a companion star. As this matter accumulates, the temperature and pressure at its base increases; eventually the accreted matter fuses into heavier isotopes. If sufficient carbon is present in the ashes of this nuclear burning, then over time the base of the ash layer can itself ignite unstably and produce an energetic explosion known as a superburst. Understanding how this layer of ash reaches sufficient temperatures to ignite carbon has been a mystery for over a decade.

A collaboration led by MSU astronomy graduate student Alex Deibel has completed a study of neutron star cooling by neutrino emission from cycles of electron captures and emissions from certain pairs of nuclei in the outer ocean of accreting neutron stars. For each pair of nuclei, there is a specific depth where these reactions occur. Strong cooling by this neutrino emission, termed Urca cooling after the Casino da Urca in Rio de Janeiro [1], can occur at depths where superbursts ignite (see the Figure) and make superburst ignition even more difficult unless additional source of heat are present.



Artist's depiction of a super burst. Credit: NASA/Dana Berry



In this schematic each dot indicates the mass number and depth of a layer in which neutrino emission from electron capture/beta decay cycles cools the neutron star crust. The size of each dot indicates the neutrino luminosity produced by each cycle; the five strongest emitters, shown in red, have a neutrino luminosity more than 25 times greater than the photon luminosity of the sun at a temperature of 0.5 GK. The shaded band indicates the depth where superburst ignition is thought to occur.

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[1] The name comes from a joke, attributed to George Gamow: that the neutrinos remove heat from the star as the casino removes money from tourists

Further Reading: A. Deibel et al. Urca Cooling Pairs in the Neutron Star Ocean and Their Effect on Superbursts. ApJ, 831: 13