Theoretical calculation of neutrino fluxes from the Sun



# Age determination

- Globular clusters: most ancient objects in the Galaxy
- Dating based on observed luminosity at "turnoff" transition and theoretical relation between luminosity and age
- depends on slowest reaction rate of CNO cycle:

## bottleneck <sup>14</sup>N(p, $\gamma$ )<sup>15</sup>O



#### History slightly biased

1949	Woodburry, Hall, Fowler
1951	Ducan, Parry
1955	Bashkin,Carlson and Nelson
1957	Lamb and Hester
1957	Pixley
1957	Hebbard and Baily
1959	Povh and Hebbard

activation activation at low energy
activation at low energy
Nal
Nal

Nal



### **History**

1949	Woodburry, Hall, Fowler	activation
1951	Ducan, Parry	activation at low energy
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first experiments on  $14N(p,\gamma)150$  in Toronto

Schröder Publication

76 87 14N(p,g)15O programm started in Münster, Bochum and Toronto targets: evaporated, implanted and gas target

high extrapolation of gs transition (Breit Wigner)



## **History**

		11101019		
1949		Woodburry, Hall, Fowler	activation	
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1957		Pixley		
1957		Hebbard and Baily		
1959		Povh and Hebbard		
1970	1973	Production of thousends of N-t production of 14N enriched	argets for 14N(α,γ)18F	
13	76	first experiments on 14N(p,γ)15 in Toronto	0	
76	87	14N(p,g)15O programm started targets: evaporated, implanted and gas target	in Münster, Bochum and	Toronto
			high extrapolation of g	s transition
		Schröder Publication	(Breit Wigner)	400
		R-matrix fit of Schröder data		E
2001		by Carmen Angulo	gs greatly reduced	10 E
2001		Tunl measurement of ANC		
0000		Tunl Doppler shift		arn
2002		measurement		
				tor (k
2002		start of LUNA2		Li fac
		start of LENA		ن E/
2003		Mukhamedzanov		0.01 <sub>E</sub>
2004		Yamada livetime		
				-
2004	today	LUNA/ LENA BOCHUM		0.001





## Status on grounstate transition



# State of the art for ${}^{14}N(p,\gamma){}^{15}O$

Transition	LUNA 2004	LENA 2005	
to	Formicola et al. Phys.Lett.B 591 (2004) 61	R.C. Runkle et al Phys. Lett. B, 94(2005)82503	
	S(0) [keV · b]		
6.79 MeV	$1.20 \pm 0.0$ ok within $\pm 2\%$ 1.15 $\pm 0.05$		
6.17 MeV	$0.08 \pm 0.03$	$0.04 \pm 0.01$	
others	$0.080 \pm 0.004$		
GS	0.25 ± 0.06	0.49 ± 0.08	
total	1.61 ± 0.0 Factor 2	$1.68 \pm 0.09$	

## <sup>14</sup>N(p,γ)<sup>15</sup>O



## Total S-factor [keVb]



# S-factor for ground state transition using AZURE by Ed Simpson



## Answer to the summing problem



Clover detector set up distance to target: 9 cm

#### Ep=380keV;Q=81,5 C



Relative to Tr =>
6.79MeV state

$$\sigma_{gs} = \sigma_{6.79} \frac{Y_{gs}}{Y_{6.79}}$$

uncertainty in  $\sigma_{dc}$  from LUNA and LENA combined 3% res contribution in 6.79 MeV  $\gamma$ -peak angular distribution in primary line but independent of

- •Charge
- Stoichometry
- ωγ

## Line shape analysis



Same as absolute But ∆ fitted No effective energy needed Check on background lines

## Result



# Set up for DS measurement



**F(**τ)=1



## Gain stability test



## Doppler shift for the 6.79MeV line



# **Doppler shift result**



Lifetime measurement of the 6792 keV state in 150 important for the astrophysical S factor extrapolation in 14N(p,)150

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Accepted in Pys Rev C

# $\Gamma_{\gamma}$ from lifetime, r-matrix and coulomb breakup



# Possible M1 component in DC

PHYSICAL REVIEW C 68, 065804 (2003)

Analyzing power measurement for the 14N(*p*, g)15O reaction at astrophysically relevant energies S. O. Nelson, 1,2 M. W. Ahmed, 1,2 B. A. Perdue, 1,2 K. Sabourov, 1,2 A. L.

Sabourov, 1, 2 A. P. Tonchev, 1, 2 R. M. Prior, 3, 2

M. Spraker, 3, 2 and H. R. Weller 1, 2

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## Remaining uncertainty



# END

- Thank you Dick for all you offered me:
- Physics
- Human relation
- Friendship
- Wisdom
- Good humor