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Animation of a neutron star X-ray burst. Credit: NASA/Dana Berr.

# First detection of the rp-process nucleus <sup>96</sup>Cd

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# Type I X-ray bursts



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Artist's conception of neutron star EXO 0748-676 (blue sphere). It is part of a binary star system, and its neighboring star (yellow-red sphere) supplies the fuel for the thermonuclear bursts.

(Image Credit: NASA)

Type I X-ray bursts are thermonuclear explosions on the surface of a neutron star accreting matter from a companion in a binary system.

#### Fuel: H and He.

Energy release: 10<sup>39</sup>-10<sup>40</sup>ergs.

Temp. range: 0.2 to 2.0 GK

Time scale: 10-100 s

Density:p ×10<sup>5</sup> g/cm<sup>3</sup>

**Mechanisms:**  $3\alpha$ ,  $\alpha$ p-process and rp-process

Recent observations from X-ray satellites

+ understanding of underlying nuclear physics

 $\Rightarrow$  constraint astrophysical models





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H. Schatz, K.E. Rehm / Nuclear Physics A 777 (2006) 601-622



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## rp-process studies at the NSCL: looking for <sup>96</sup>Cd



Primary beam: <sup>124</sup>Xe at 140 MeV/u. Courtesy: R. Fontus "L^J lon K500 source K1200 **S800** <sup>96</sup>Cd cocktail Be target 188mg/cm<sup>2</sup>. sent to vault Al wedge 180 mg/cm<sup>2</sup> A1900 settings: Bρ<sub>1,2</sub>=3.1675 T·m; Bρ<sub>3,4</sub>=2.7965 T·m Momentum acceptance = 1%

Average beam current = 5.6 pnA





## Experimental challenges in rp-process studies at NSCL



First observation of <sup>60</sup>Ge and <sup>64</sup>Se Stolz, et. al. Phys. Let. B. 627 (2005) 32.

- The existence of <sup>60</sup>Ge and <sup>64</sup>Se was demonstrated for the first time.
- Difficulties:

Lack of beam purity limited the total beam intensity to about 1/10 of what would have been available.

Total beam intensity ~  $10^3 \text{ pps}$ , did not allow  $\beta$ -decay measurements.

Proper correlation of events imposes a limit on the implant rate.

Example: 96Cd

Estimated half-life = 1s

Assuming beam spot covering ~1200 pixels

Wait 5 half-

 $\Rightarrow$  Average



# We need to add a velocity filter to the A1900 configuration





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## rp-process studies at the NSCL: looking for <sup>96</sup>Cd



#### Primary beam: <sup>124</sup>Xe at **Experimental station** 140 MeV/u. Courtesy: R. Fontus "L^J lon K500 source K1200 **S800** <sup>96</sup>Cd cocktail **RFFS:** Additional Be target 188mg/cm<sup>2</sup>. sent to vault purification of secondary proton-rich beams Al wedge 180 mg/cm<sup>2</sup> A1900 settings: Bρ<sub>1,2</sub>=3.1675 T·m; Bρ<sub>3,4</sub>=2.7965 T·m

Momentum acceptance = 1 %

Average beam current = 5.6 pnA





## The Solution: The Radio Frequency Fragment Separator







## RFFS simulated in LISE++



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Example of beam purification using RFFS. Settings on <sup>96</sup>Cd.





#### RFFS Commissioning Run Experimental Setup MICHIGAN STATE MICHIGA

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## <sup>96</sup>Cd: Particle Identification

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## <sup>96</sup>Cd: Particle Identification













## Final Remarks



The proton rich N=Z=48 nucleus <sup>96</sup>Cd has been observed for the first time and it has a half-life longer than 95 ns. Our measurement therefore paves the way for direct studies of the decay of this nucleus, which will constitute important pieces of information for the rp-process models.

> The efficiency of this measurement was greatly improved by the additional purification of the cocktail beam provided by the RFFS.

>Three experiments that will make use of the RFFS have been approved by the NSCL PAC31:

- Beta-decay Half-life Measurement of <sup>84</sup>Mo
- Beta-Delayed Proton Emission of <sup>69</sup>Kr
- Study of the Beta-Decay of <sup>100</sup>Sn

>The measurement of production rates of nuclei far from stability with existing facilities is crucial for the prediction of rates achievable with future rare isotope acceleration facilities.





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