

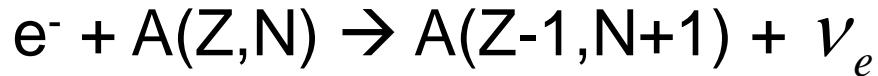
Comparison Between Measured Gamow-Teller Distributions and the Corresponding Electron Capture Rates for pf-shell Nuclei in Pre-supernova Stars



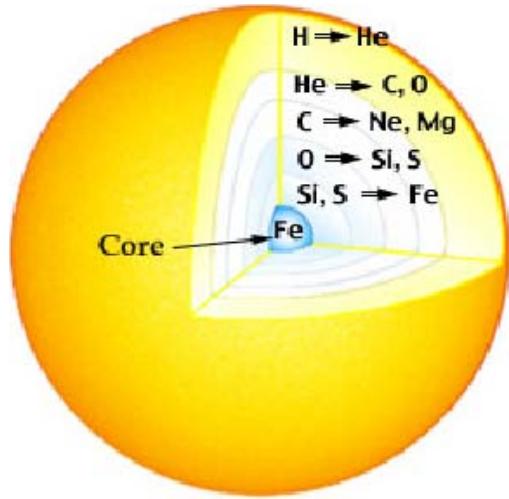
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Electron Capture and Supernovae

Electron Capture:



Core-Collapse Supernova



<http://www.solstation.com/x-objects/xte-bh.htm>

Type Ia Supernova



David A.Hardy
www.astroart.org & PPARC

Influences Dynamics
of Supernova

Provide constraints
on models

Electron Capture and B(GT)

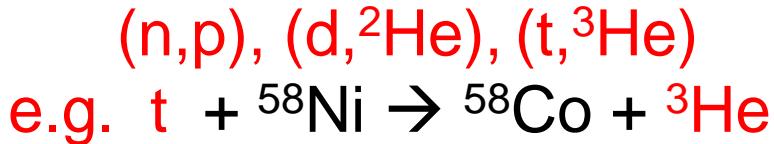
Electron Capture \Rightarrow Gamow-Teller Transitions:

Gamow-Teller (GT)
(spin-flip)

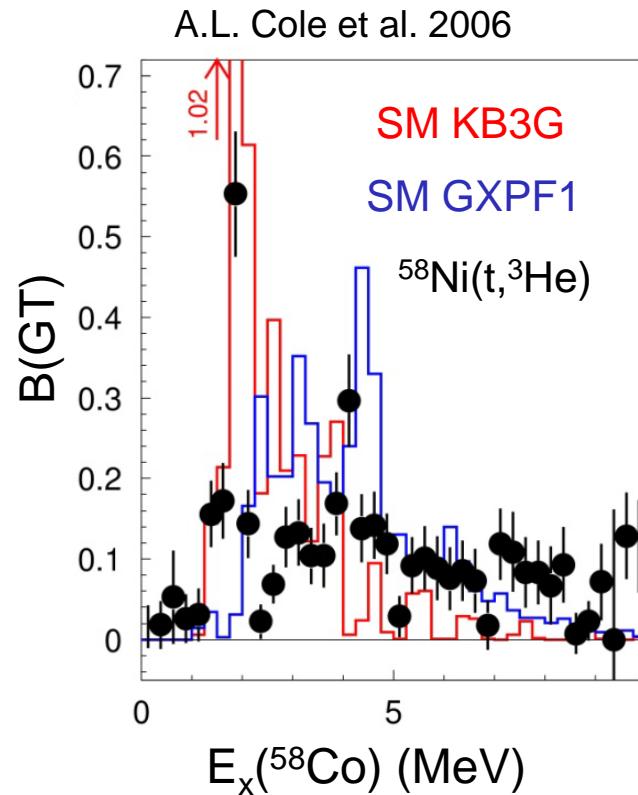
$$\Delta S = 1, \Delta L = 0, \Delta J = 1, \Delta T = 1$$

Transition Strength = B(GT)

B(GT) measured via
charge-exchange



$$B(GT) \propto \frac{d\sigma(q=0)}{d\Omega}$$



Electron Capture Rate \propto B(GT)

Core-Collapse Supernova and B(GT)

Electron Capture Rate \propto B(GT)

Consider two models for calculating B(GT)

WW (Woosley, Weaver)

Fuller, Fowler, Newman (FFN) Weak Interaction rates

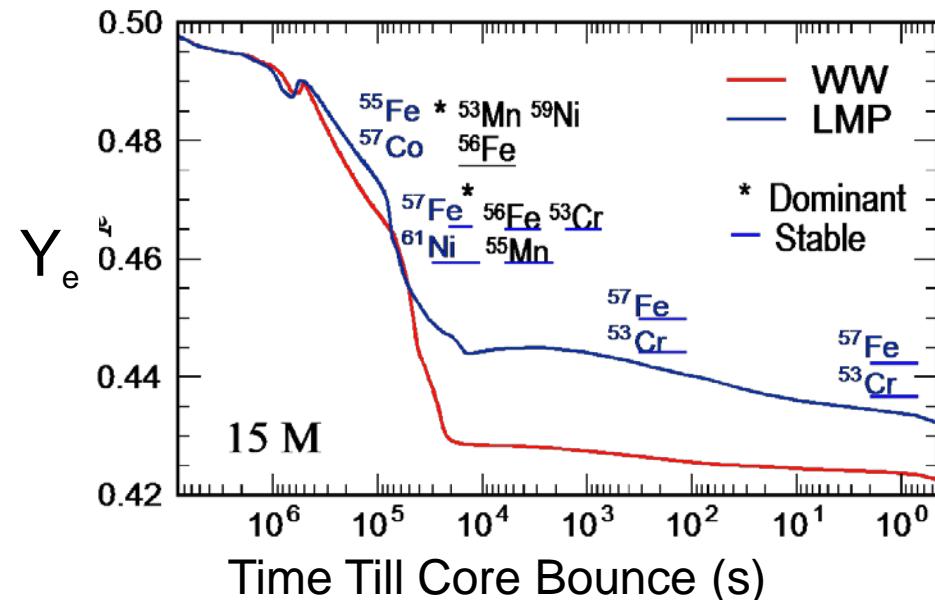
- Independent Particle Model (IPM)
- No interaction b/t valance nucleons

LMP (Langanke, Martinez-Pinedo)

& Heger, Langanke, LMP Weak Interaction rates

- Shell Model (SM) Calculations
- Interaction b/t valance nucleons leads to fragmentation and quenching of B(GT)

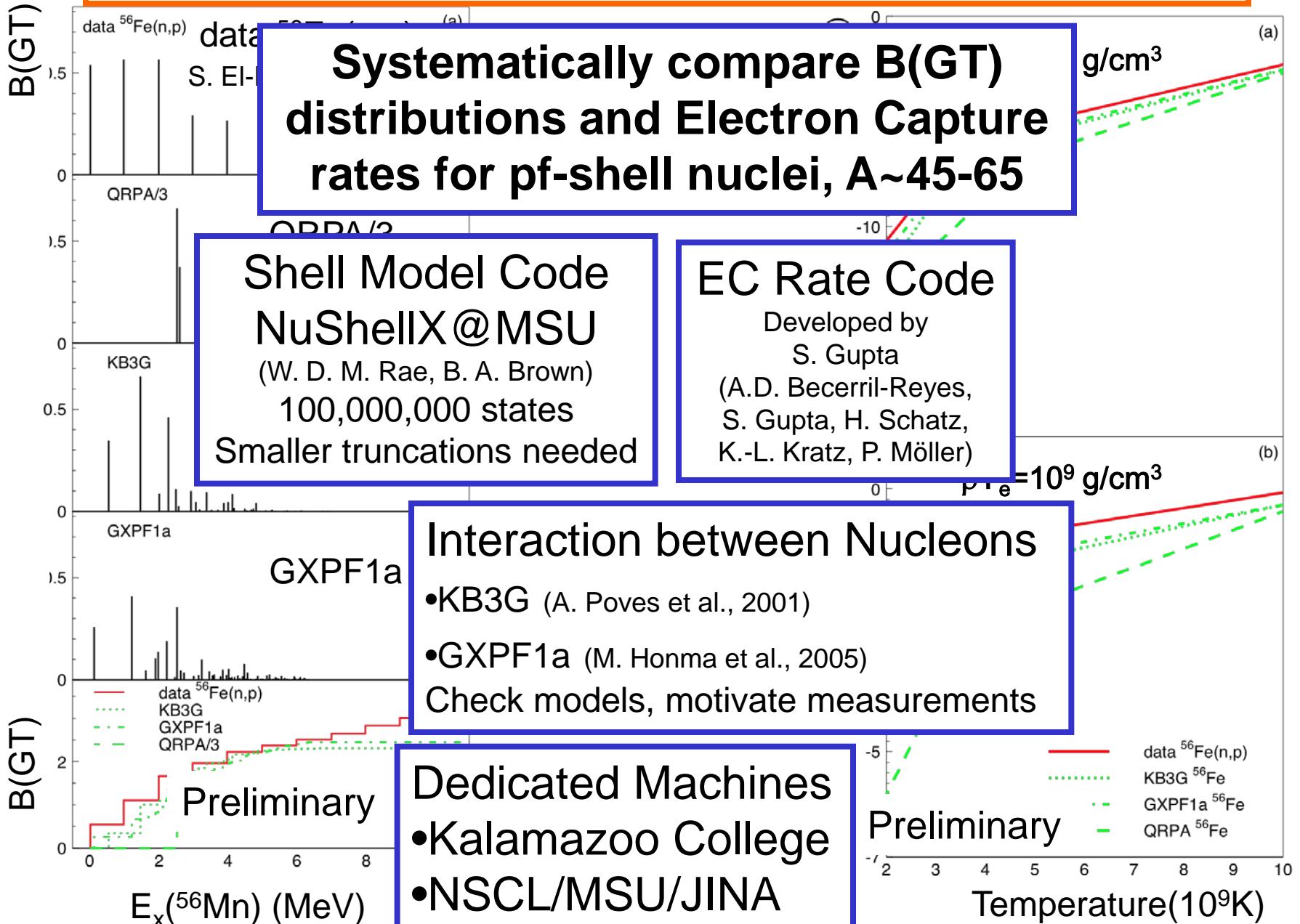
Heger et al. (2001)



- pf-shell nuclei A~45-65 important
- Stable and Radioactive Nuclei
- Several Nuclei are important

Can Not Measure them all

Calculate and Compare



Where to Start?

Nuclei for which there exist measured $B(GT)$:

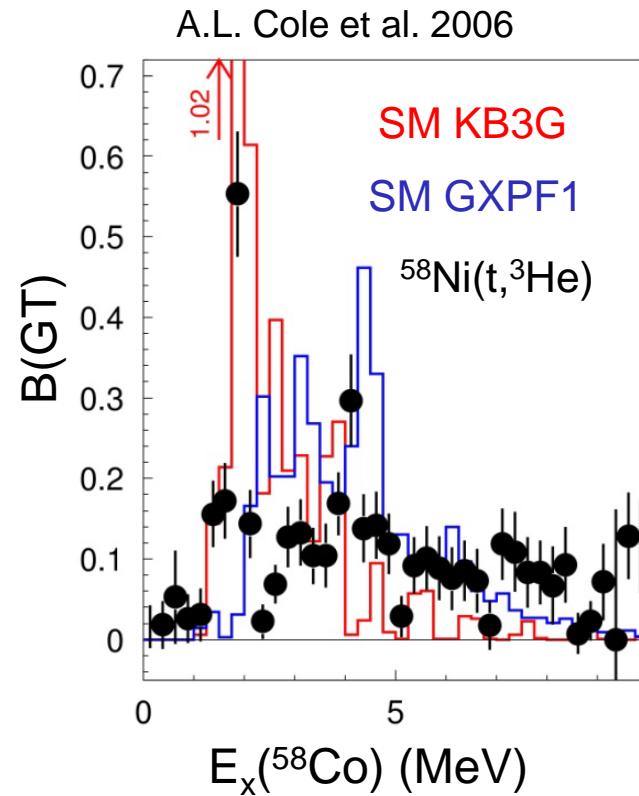
^{45}Sc , ^{48}Ti , ^{50}V , ^{51}V , ^{55}Mn , ^{54}Fe , ^{56}Fe , ^{59}Co , ^{58}Ni , ^{60}Ni , ^{62}Ni , ^{64}Ni , ^{64}Zn

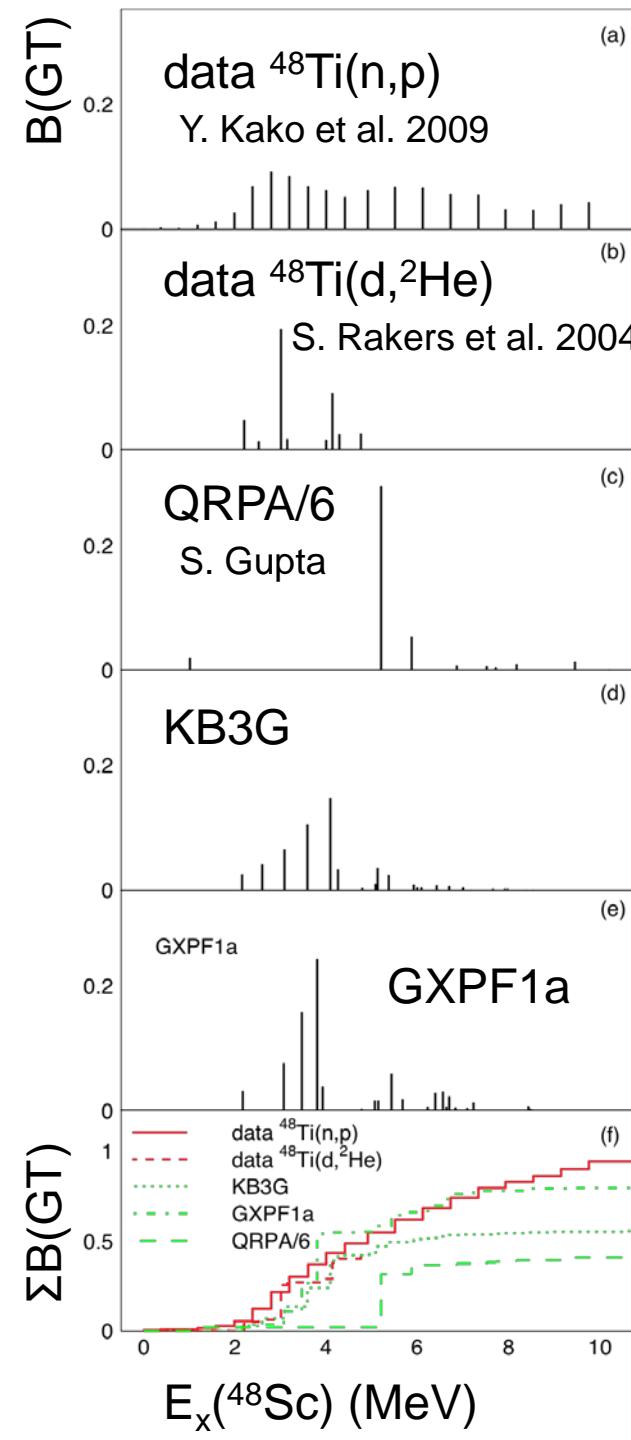
**$B(GT)$ measured via
charge-exchange**

(n,p), (d, ^2He), (t, ^3He)

e.g. t + $^{58}\text{Ni} \rightarrow ^{58}\text{Co} + ^3\text{He}$

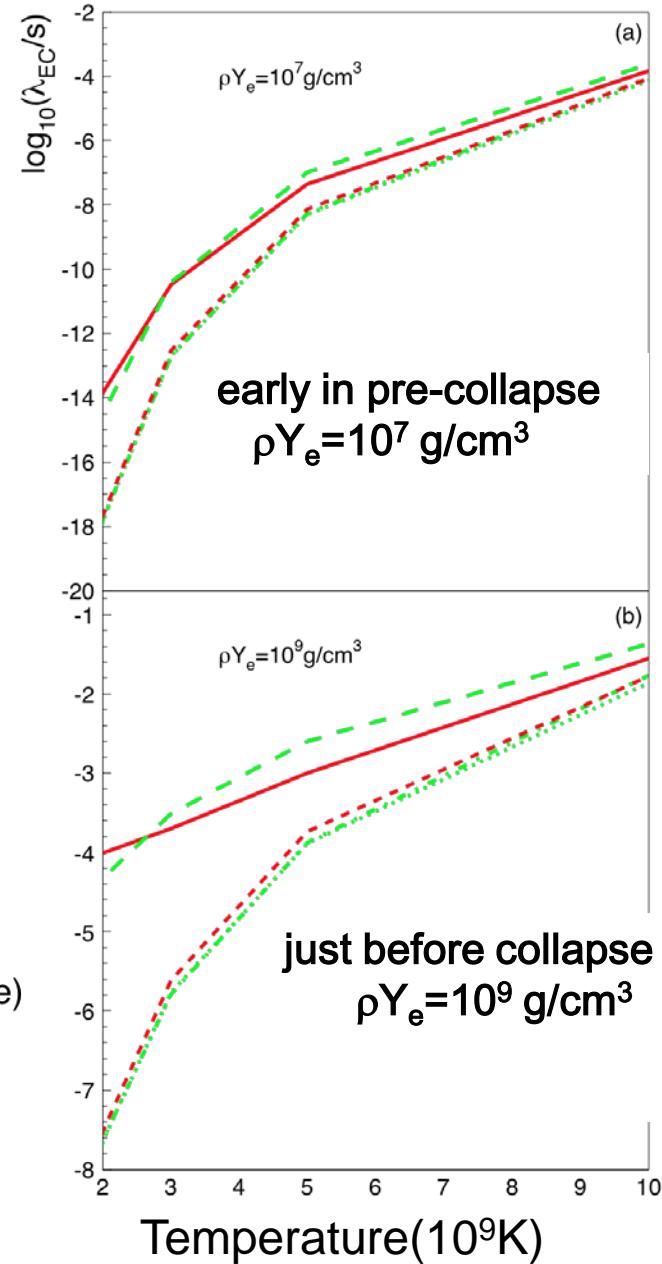
$$B(GT) \propto \frac{d\sigma(q=0)}{d\Omega}$$



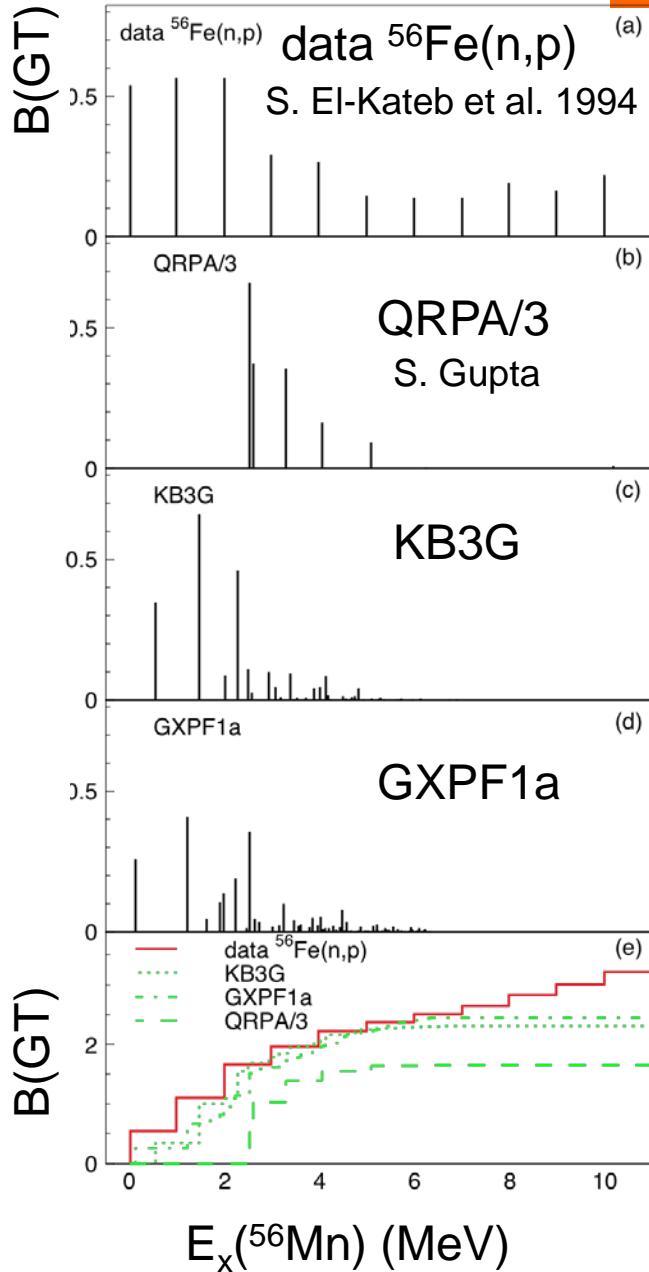


Preliminary

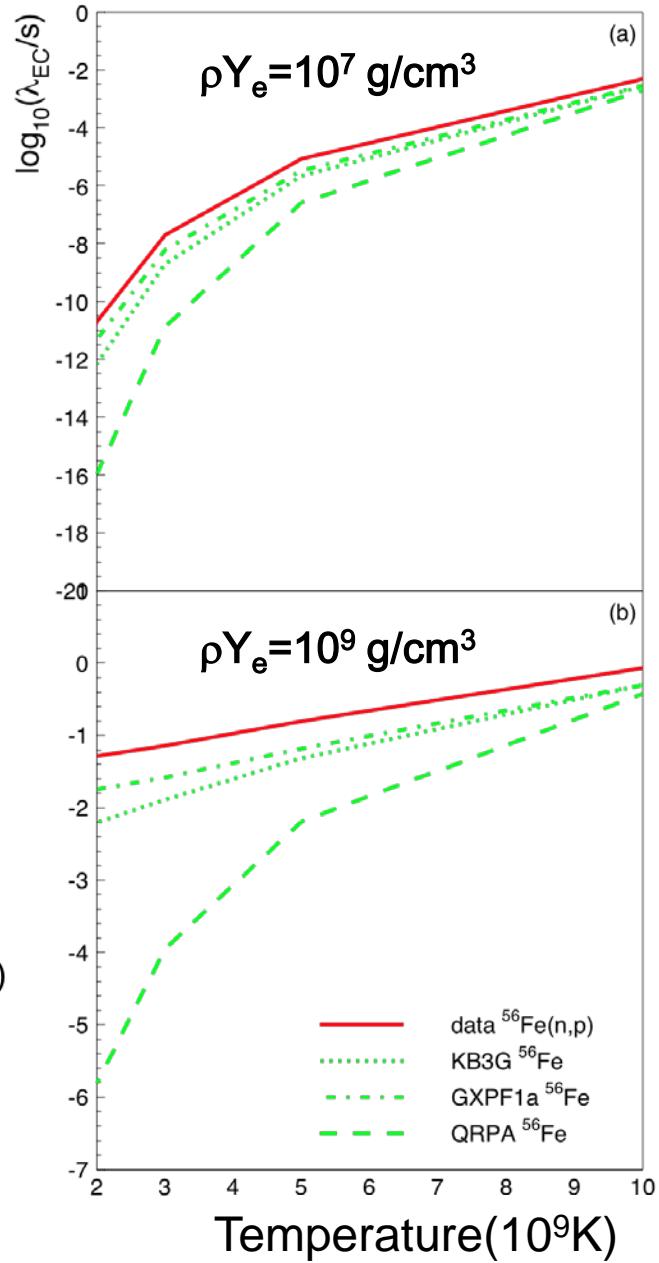
$\lambda_{\text{ec}} \propto B(\text{GT}),$
state energy
and strength



$^{56}\text{Fe} \rightarrow ^{56}\text{Mn}$



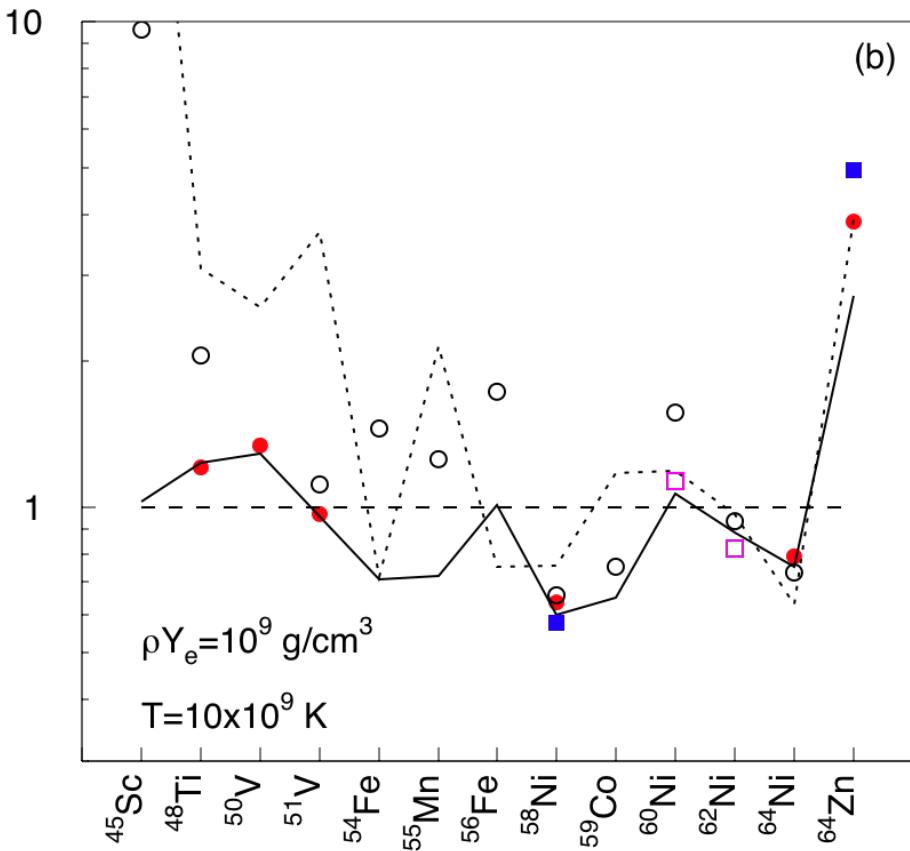
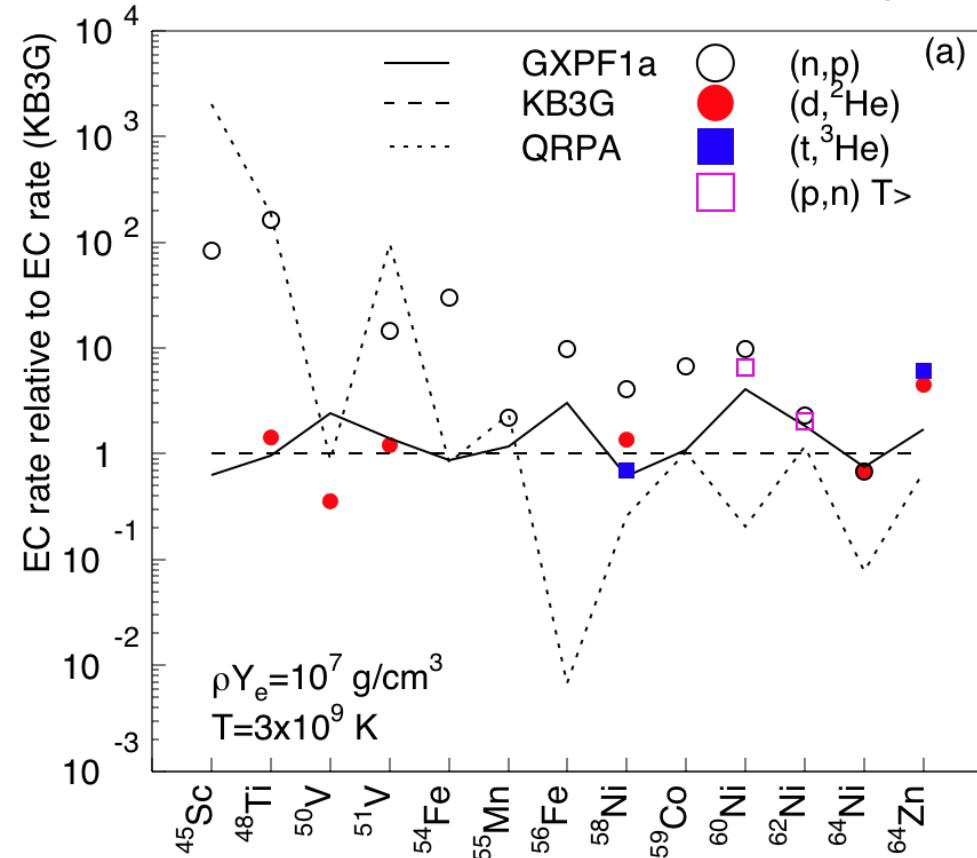
Preliminary



$\lambda_{\text{ec}} \quad ^{56}\text{Fe(gs)} \rightarrow ^{56}\text{Fe (1/s)}$

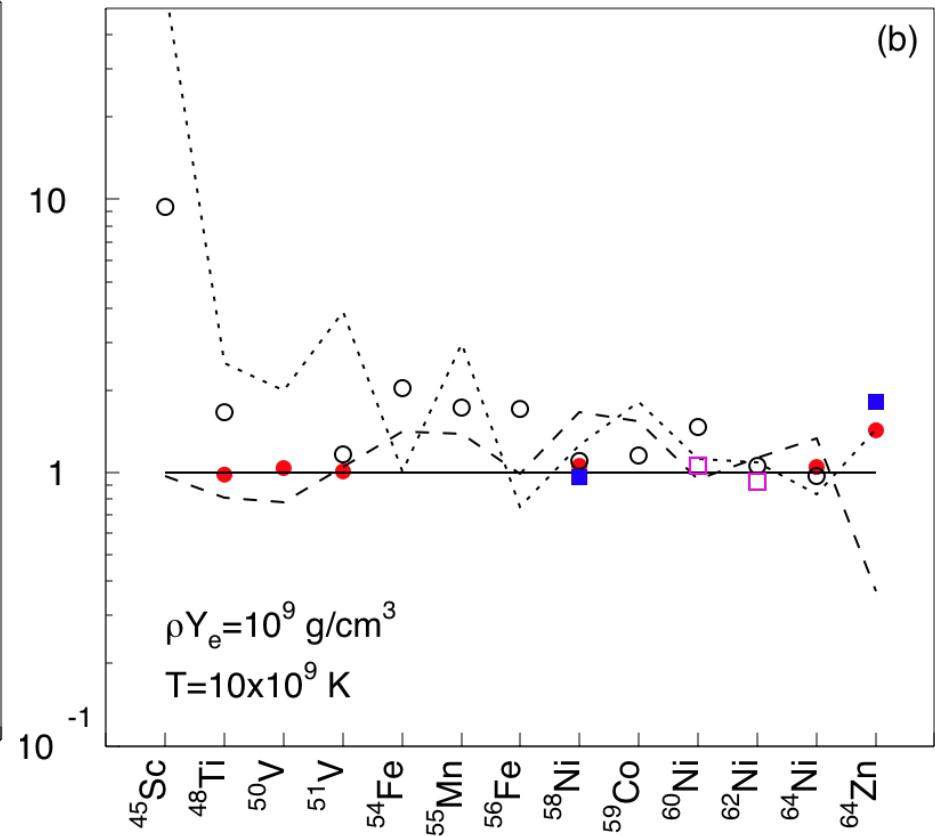
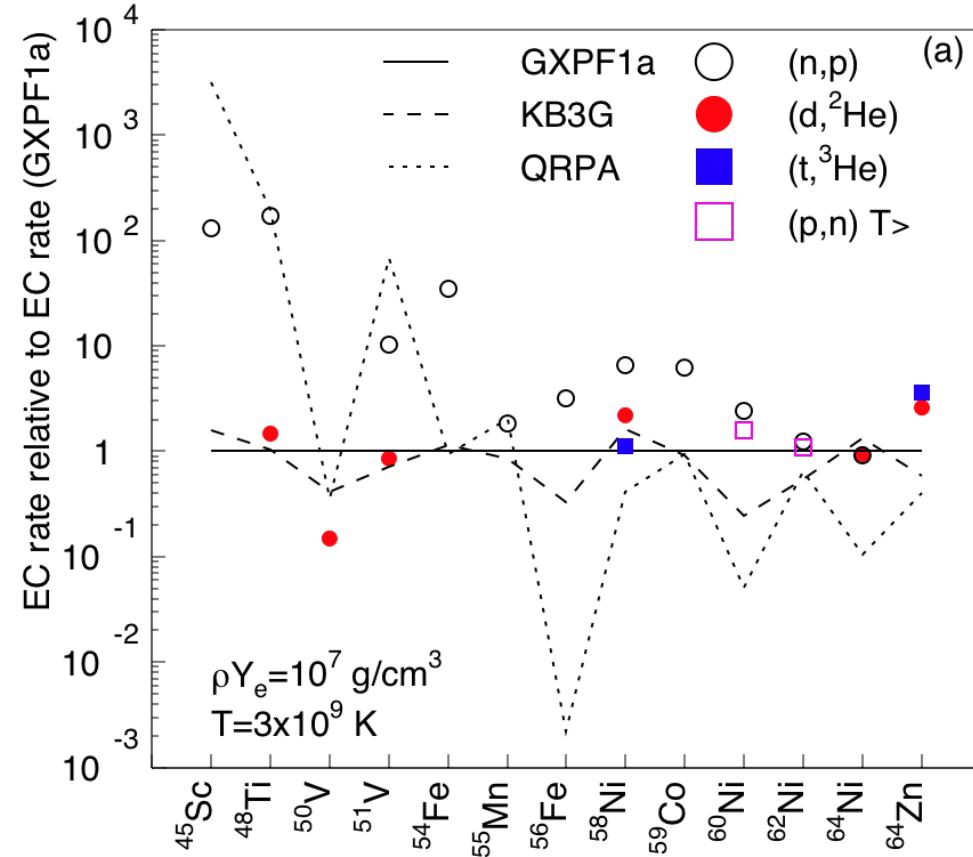
EC Rate Comparison: KB3G

Preliminary



EC Rate Comparison: GXPF1a

Preliminary



Collaboration

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