Online Tools for Understanding the Equation of State

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- Condition of complete equilibrium among nuclear species under exchange of neutrons and protons
- Achieved in late stages of stellar evolution or in supernova explosions

$$\mu(Z,A) = Z\mu_p + (A-Z)\mu_n$$

- Our assumptions about the nuclei:
 - Boltzmann particles (except at low temperature where they can be bosons)
 - Incompressible, but finite volume
- Our assumptions about the nucleons (and electrons and positrons):
 - Non-interacting
 - Fully relativistic and degenerate

- The nuclei: $\mu(Z,A) = m(Z,A)c^{2} + kT \ln\left(\frac{\rho N_{A}Y(Z,A)}{G(Z,A)} \left[\frac{h^{2}}{2\pi m(Z,A)kT}\right]^{3/2}\right)$
- We thus require:
 - Masses: m(Z,A)
 - Ground state spins: J(Z,A)
 - Nuclear partition functions: G(Z,A), which we take to be a multiple of 2J(Z,A)+1

- NSE codes are fairly straightforward to write and run
- However, one must write or obtain the code, compile it, test it, and then data mine the output
- This may be a barrier to those who would like to use such codes but do not have the time to undertake these steps

http://nucleo.ces.clemson.edu/pages/nse

- Online tool for calculating NSE
 - User inputs in a straightforward way nuclear properties (masses, spins, partition functions) or uses default values.
 - Computer at Clemson calculates NSE for the user-input set of temperatures, densities, and electron fractions.
 - Interfaces allow user to explore results of the calculation, including abundances and thermodynamic properties of the matter

We have chosen W3C standards or recommendations

- Input data upload from client to server is via XML (eXtensible Markup Language)
 - Validation via Schema checking
 - Formatting via XSLT
- Form: <tag>data</tag>

Example XML file

<nuclear_data>

<!-- Neutron -->

<nuclide> <z>0</z> <a>1 <mass_excess units="MeV">8.071</mass_excess> <spin>0.5</spin> </nuclide>

<!- Proton -- >

</nuclide> <z>1</z> <a>1 <mass_excess> 7.289 </mass_excess> <spin>0.5</spin> </nuclide>

</nuclear_data>

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Formatted Input

Ζ	Α	Mass Excess (MeV)	Spin
0	1	8.07100	0.500000
1	1	7.28900	0.500000
1	2	13.1360	1.00000
1	3	14.9500	0.500000
2	3	14.9310	0.500000
2	4	2.42400	0.00000
3	6	14.0850	1.00000
3	7	14.9070	1.50000
3	8	20.9450	2.00000

Output

- Output format is binary or FITS file
- User may data mine the output as soon as it is available or download it. Downloaded output files may be uploaded again.
- Data mining and graphics are through IDL on the Net (ION).

Calculations

- Single calculations: user inputs T, density, Y_e
 These calculations usually take < 1 minute
- Multiple calculations: user inputs range of T, density, or Y_e or enters an NSE trace
 - These calculations can take a while
 - The computer puts the calculation in the background and then emails the user with information on how to access the data when the calculation is done

Example calculation

- T = 1 MeV
- $Y_e = 0.4$
- Density = 1.0e-6 to 1 fm^{-3}













Mass Fraction





Pages available

- Ideal gas calculator
 - Computes thermodynamics of an ideal fermionic or bosonic gas for any degree of degeneracy or relativity
 - Available:
 - http://nucleo.ces.clemson.edu/pages/ideal_gas/0.1
- NSE calculator
 - Computes abundances and thermodynamics of matter in nuclear statistical equilibrium
 - Available:
 - http://nucleo.ces.clemson.edu/pages/nse/0.1
- Galactic Chemical Evolution calculator
 - Computes simple models of one-zone galactic chemical evolution
 - Available:
 - http://nucleo.ces.clemson.edu/pages/cugce/

Purposes

- Research
- Education
- Code Archiving

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