

(n, γ) experiments and their implications to nuclear astrophysics

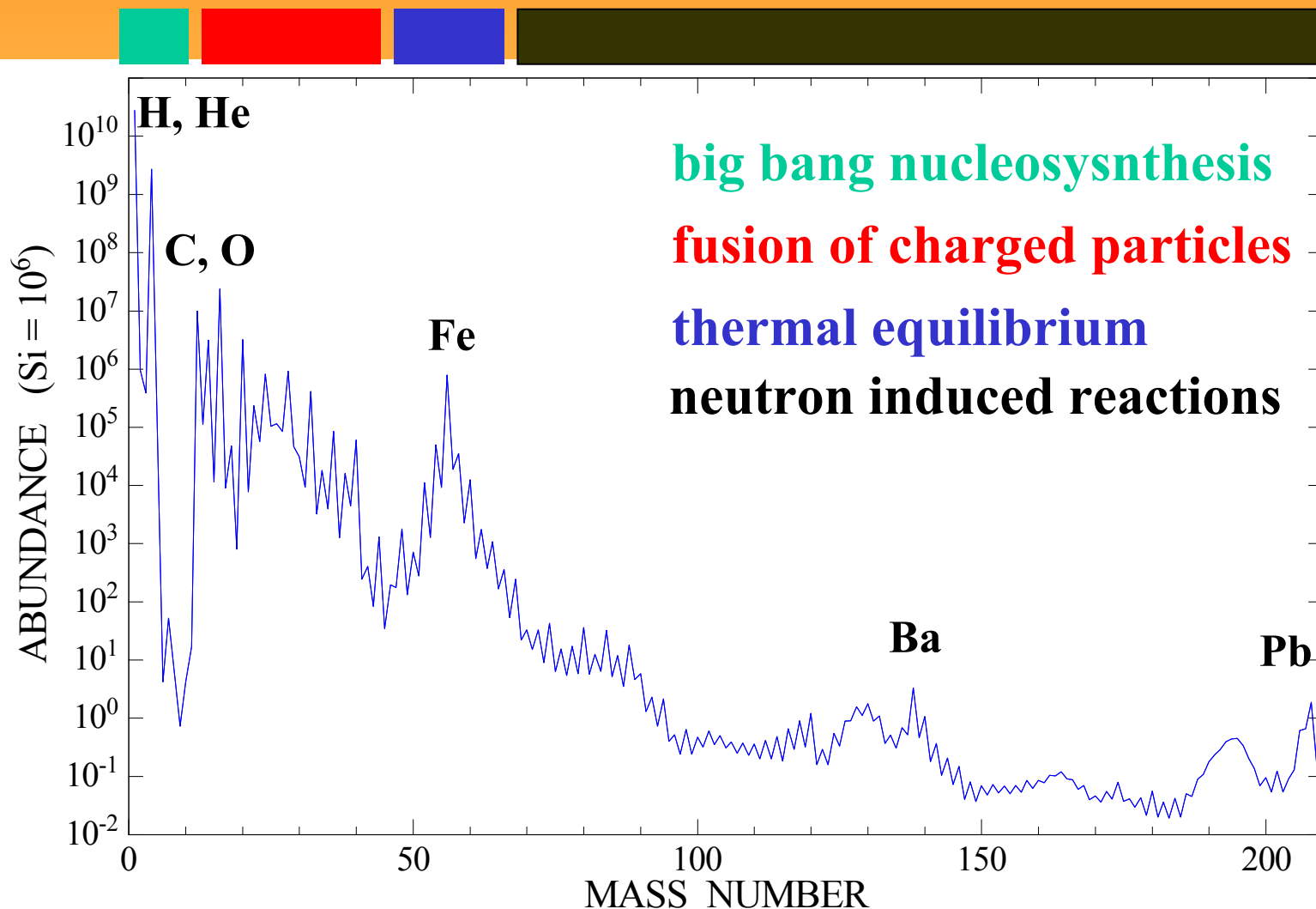
René Reifarth

University of Notre Dame - Nuclear Physics Colloquium
Notre Dame, Indiana, January 26, 2004

Outline

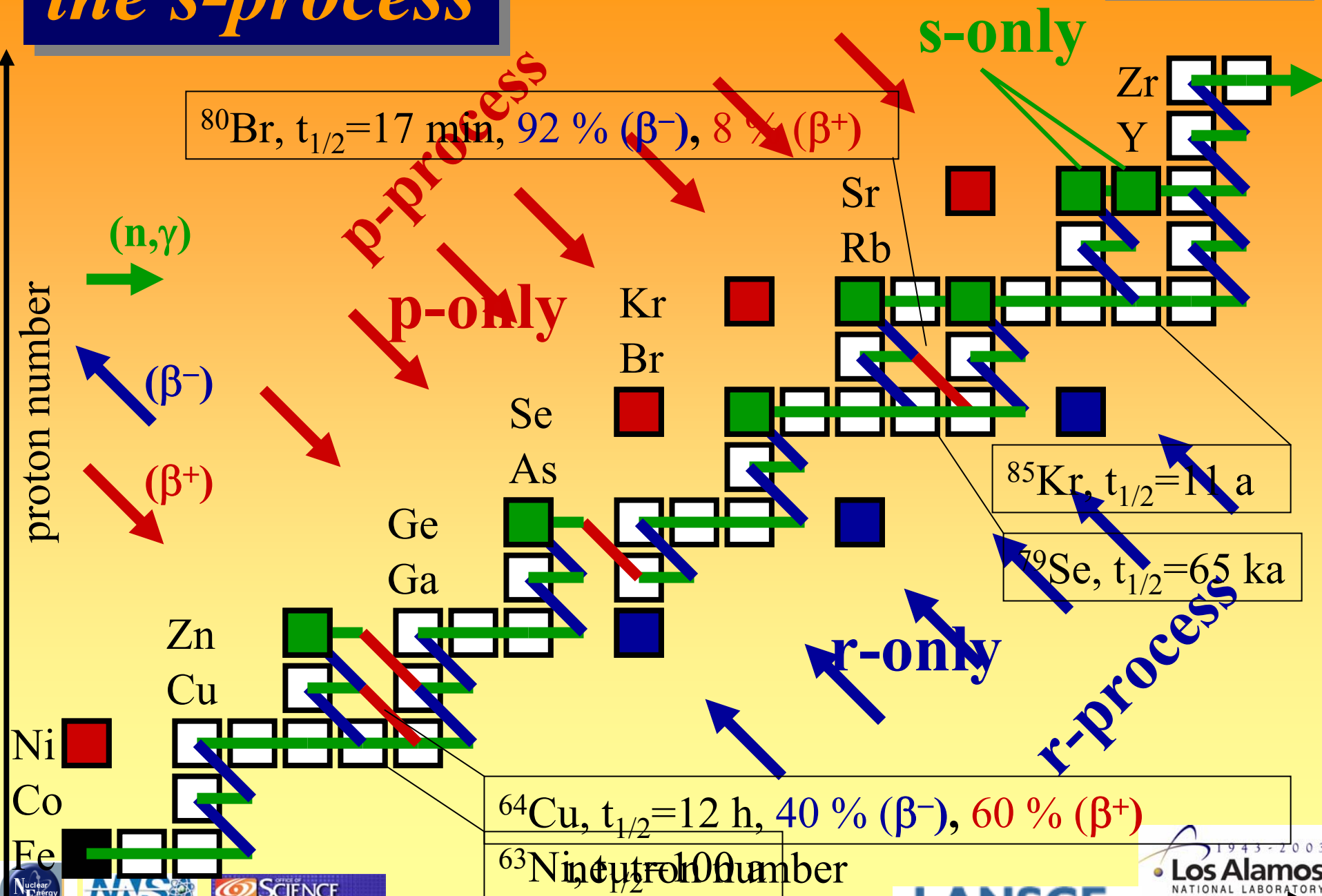
- Astrophysics – s-process
- (n, γ) experiments
 - Activation technique
 - Time Of Flight technique
- Outlook

solar abundance distribution



the s-process

Astrophysics



s-process - unbranched

conservation of nuclei: cross section x abundance = constant

$$\left(\frac{dN_A}{dt} \right)^{\text{destruction}} = \left(\frac{dN_{A+1}}{dt} \right)^{\text{production}} \approx \left(\frac{dN_{A+1}}{dt} \right)^{\text{destruction}}$$

s-process path
→

A

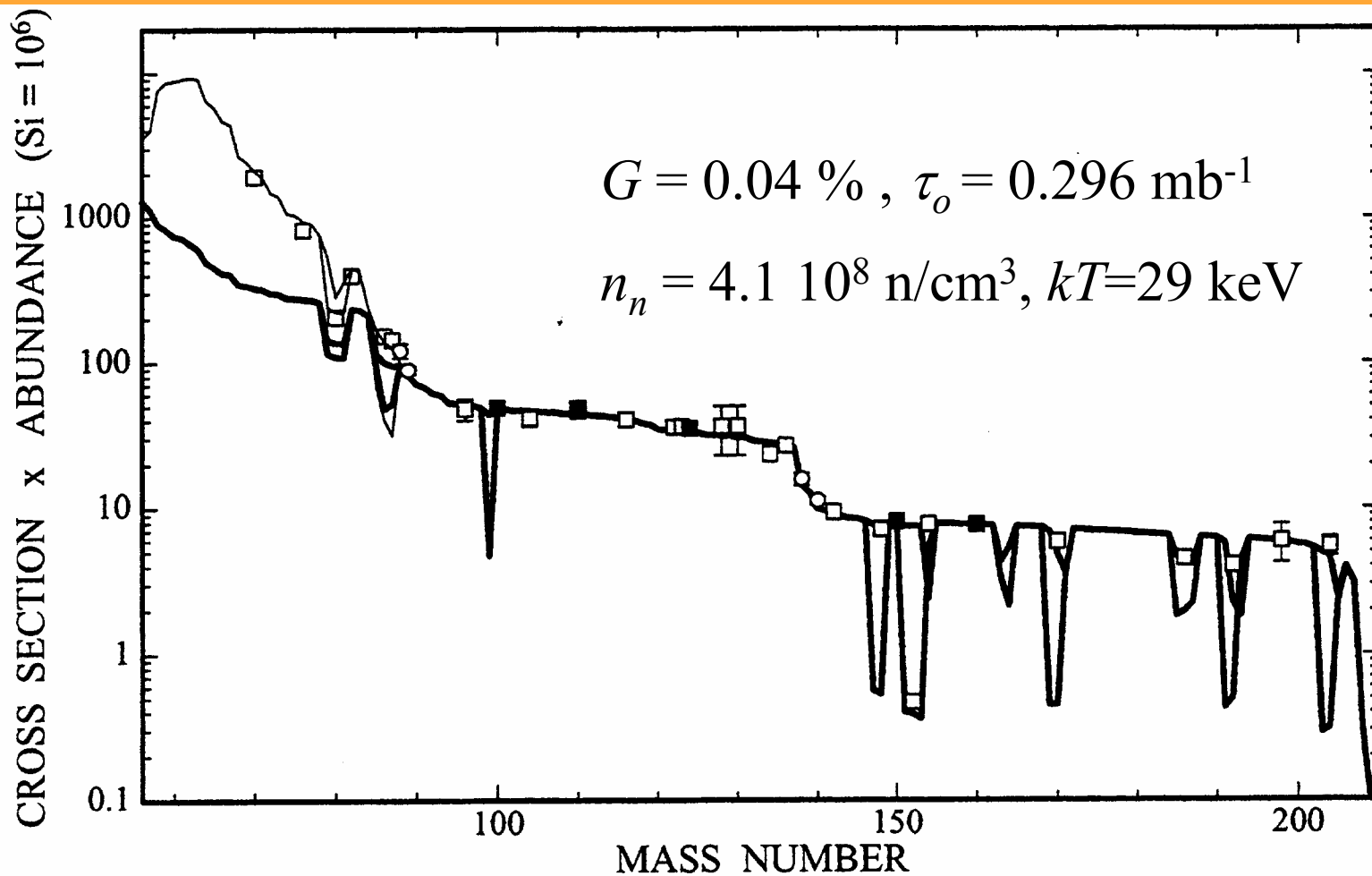


A+1

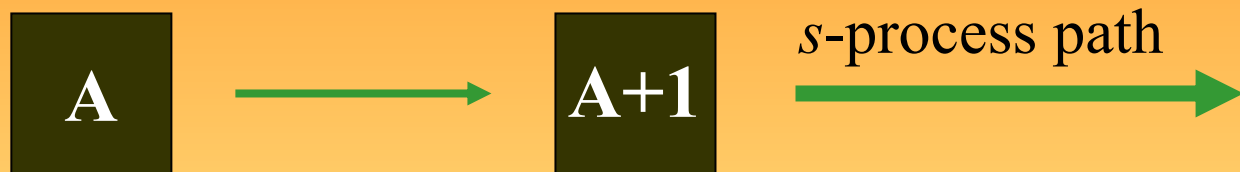
s-process path
→

$$\langle \sigma_A \rangle N_A \approx \langle \sigma_{A+1} \rangle N_{A+1}$$

$N\sigma$ -systematics from the Classical Model

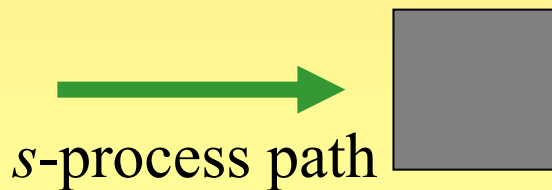


s-process – branched



$$\lambda_{\beta} \sim t_{1/2}^{-1}$$

A diagram showing a branching point. A blue arrow points from a box labeled **A+1** down and to the left to a box labeled **A**. A green arrow points from a box labeled **A** to a box labeled **A+1**. A green arrow points from a box labeled **A** to a box labeled **A+1**. A green arrow points from a box labeled **A** to a box labeled **A+1**.



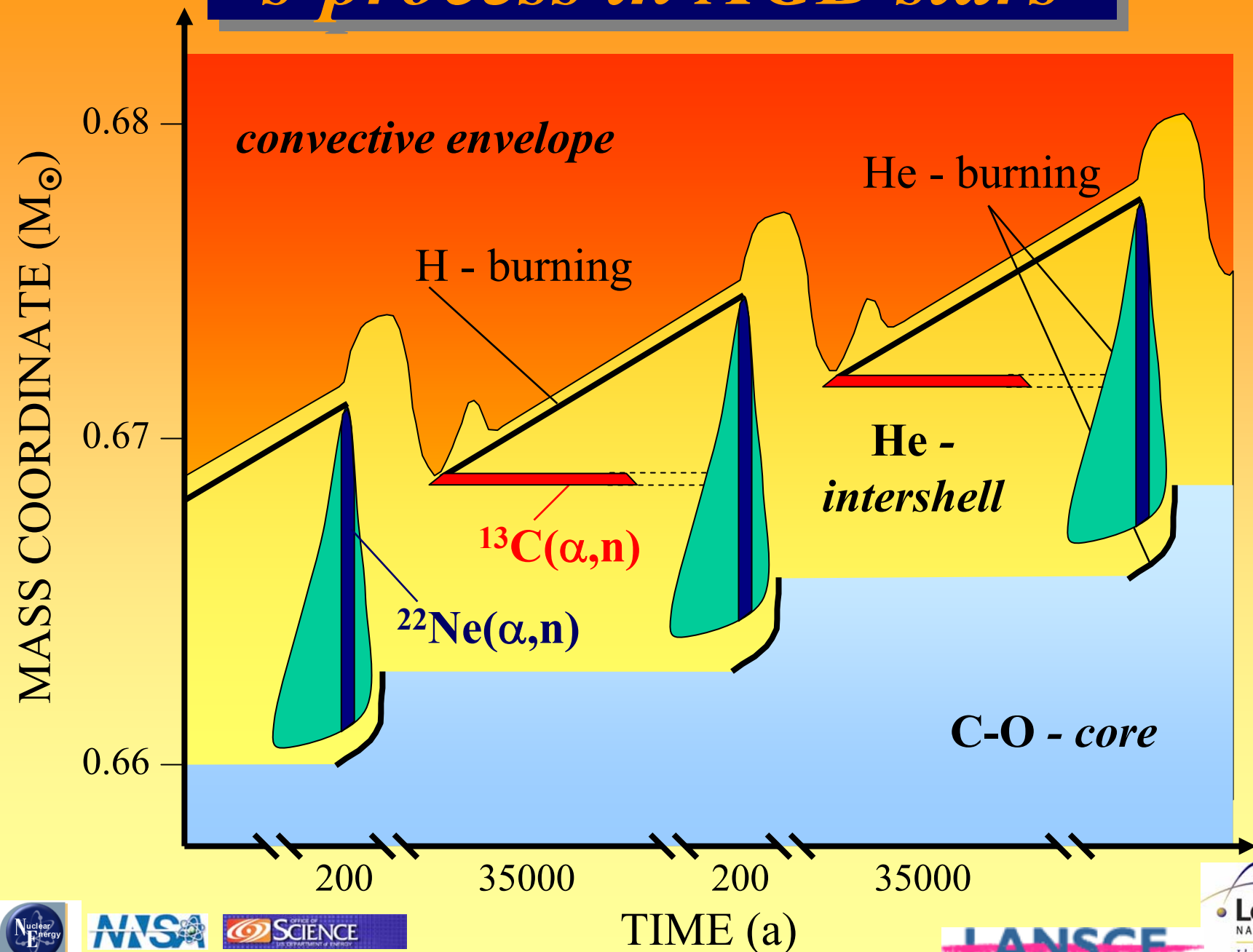
$$\lambda_n \sim \sigma_{\text{capture}} \phi_n$$

$$f_{\beta} = \frac{\lambda_{\beta}}{\lambda_{\beta} + \lambda_n}$$

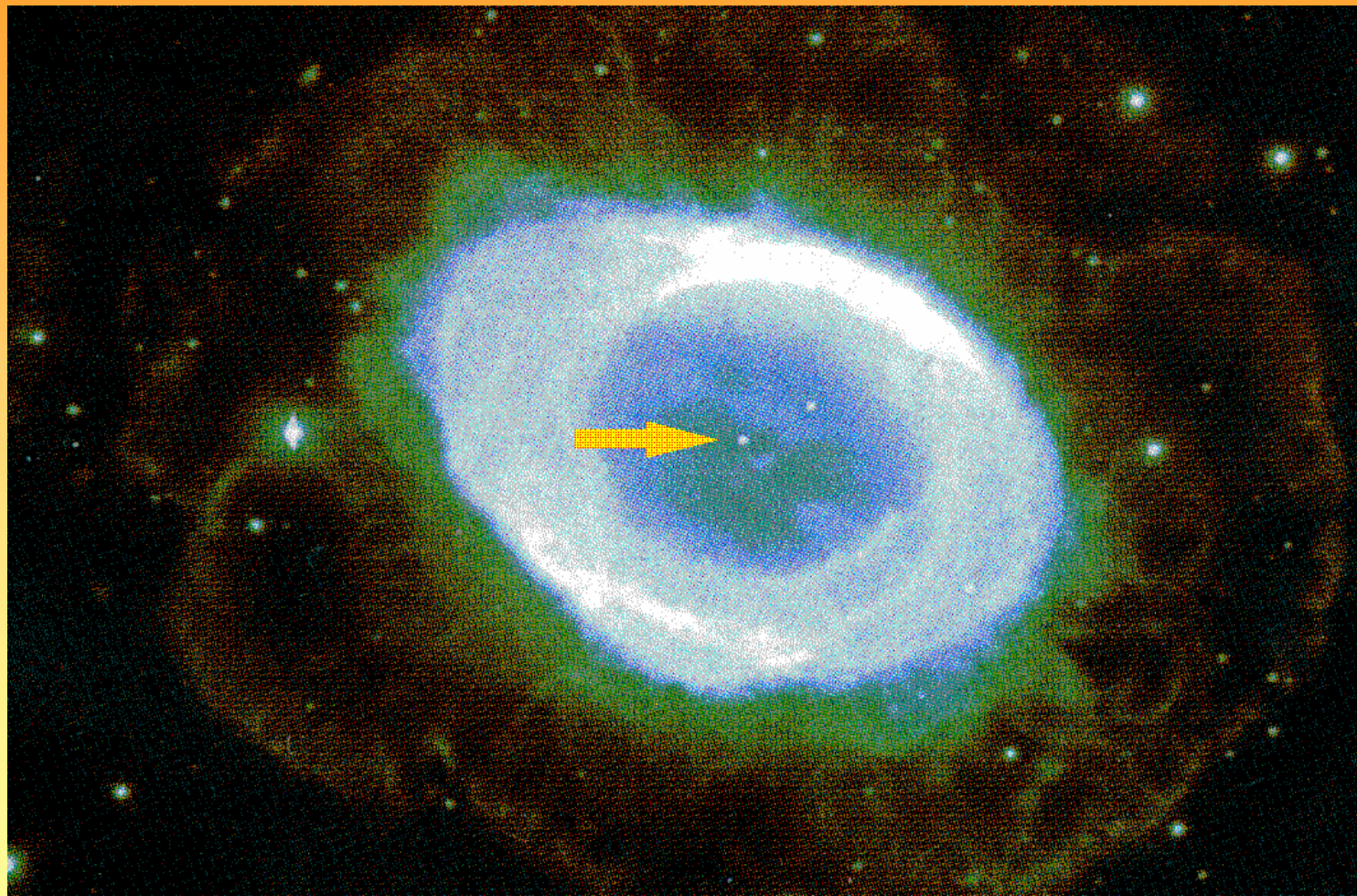
$$\approx \frac{\langle \sigma \rangle_A N_A}{\langle \sigma \rangle_{A+1} N_{A+1}}$$

s-process in AGB stars

Astrophysics



The stellar site

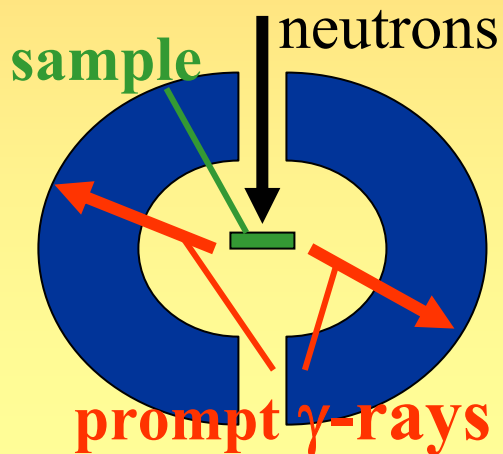


(n, γ) @ radioactive isotopes

Experiments

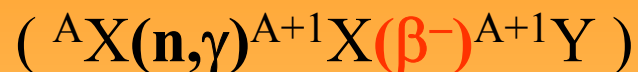
PROMPT γ -detection (4π - scintillators)

- 👍 TOF experiments
- 👍 intense (white) neutron sources
- 👎 sensitivity to background

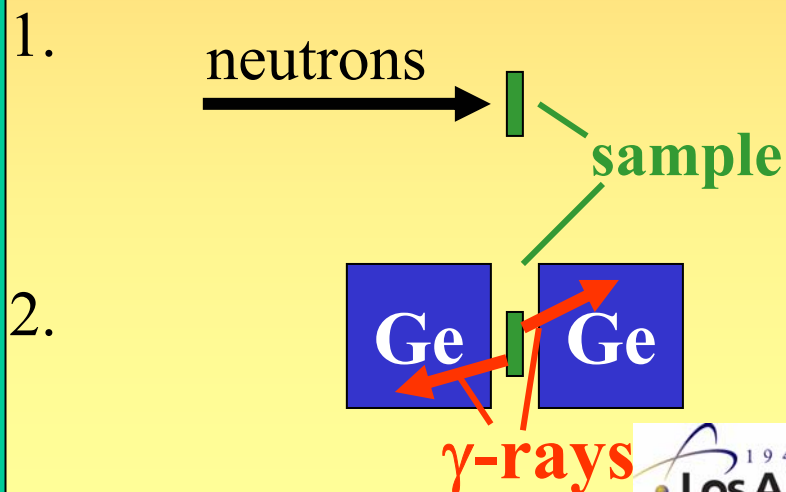


Activation technique -

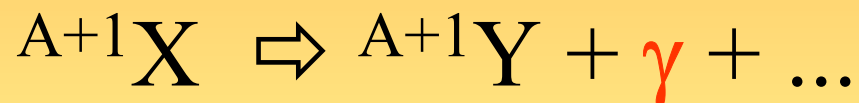
DELAYED



- 👍 very high sensitivity ($\sim \mu\text{g}$)
- 👎 only average neutron energies
- 👎 only if ${}^{A+1}X$ is “reasonable” radioactive



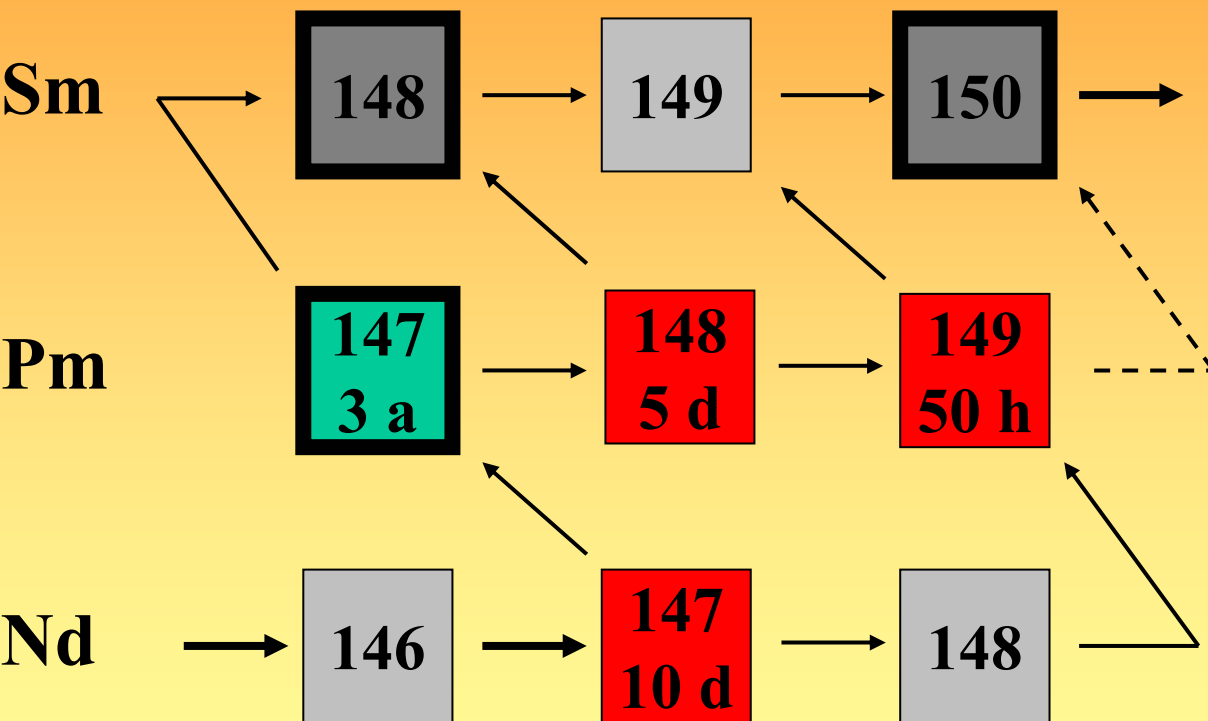
Evidence for neutron capture: **ACTIVATION**



Produced Activity:

$$A \propto \frac{{}^AN \cdot \Phi_n \cdot \sigma \cdot t_a}{t_{1/2}}$$

Branch point at ^{147}Pm



- 2 *s-only* isotopes
- *neutron flux* dependent
- almost independent from temperature, electron density

⇒ *stellar neutron monitor*

Samples

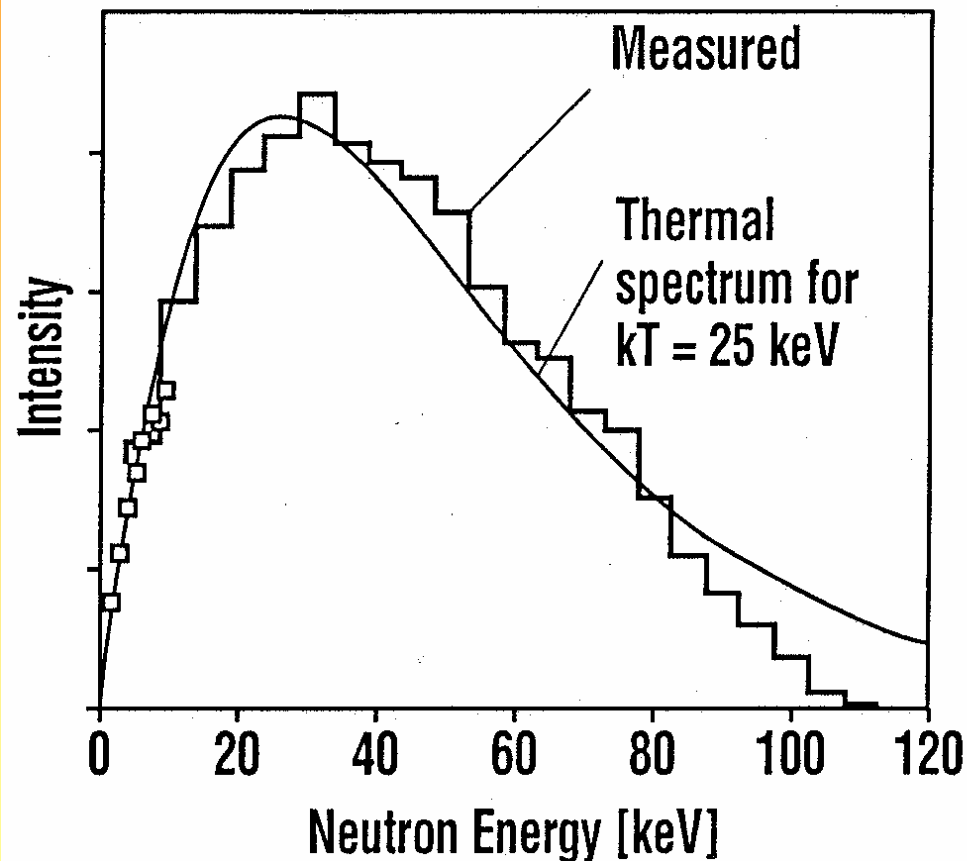
- 28 ng ¹⁴⁷Pm ($t_{1/2} = 2.6$ y, 962 kBq)
- carrier: isotopically enriched ¹⁴⁷Sm
- canned in 0.3 mm graphite
- impurities: 0.0015 % ¹⁴⁶Pm ($t_{1/2} = 5.5$ y)

Neutron spectrum

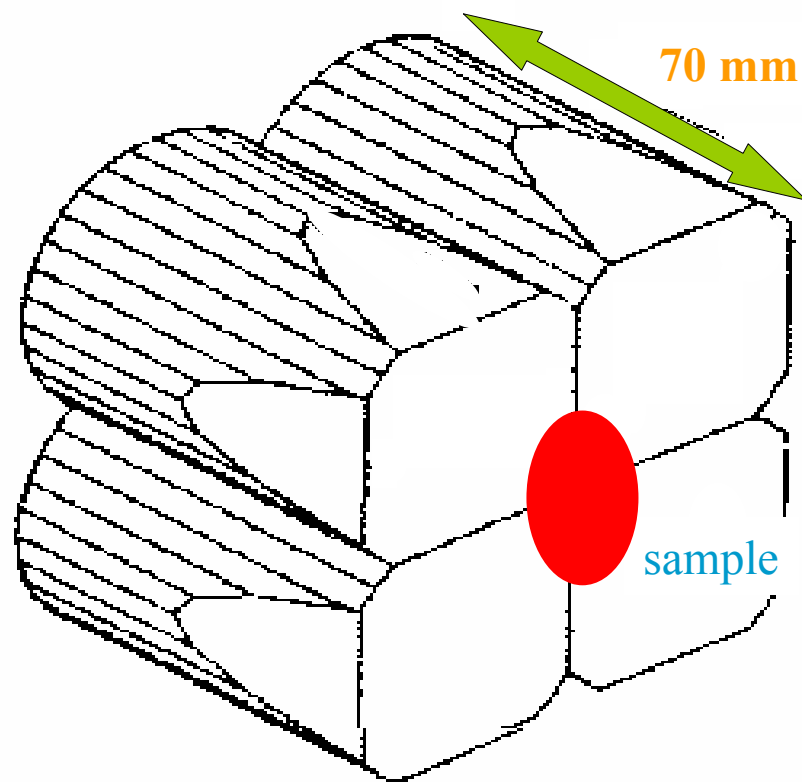
Quasi-Maxwellian
averaged distribution:

$$kT = 25 \text{ keV}$$

$$E_{max} = 110 \text{ keV}$$



γ -detection



2 Ge-Clovers, face to face

Efficiency @ 1115 keV:

single crystal:

$$\epsilon_{\text{tot}} = 11 \%$$

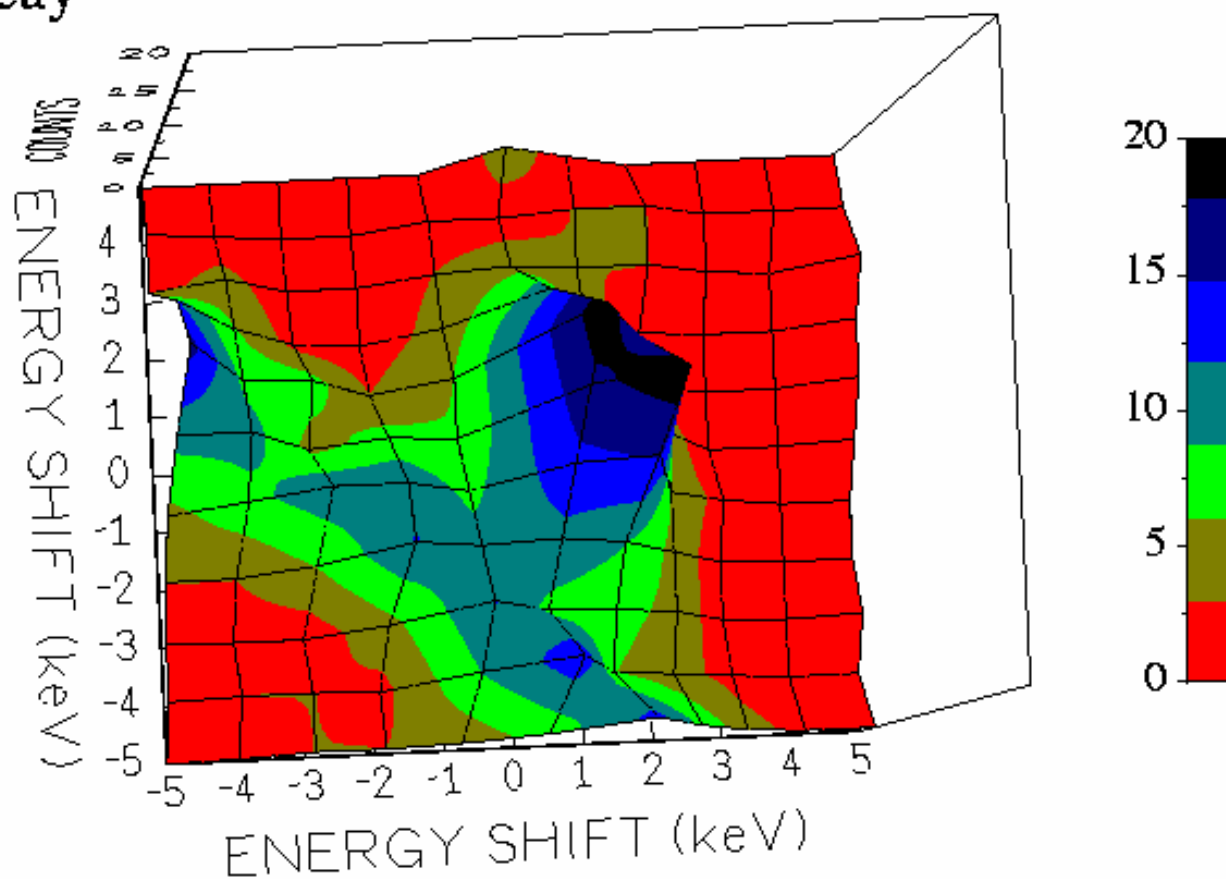
$$\epsilon_{\text{peak}} = 1.1 \%$$

addback:

$$\epsilon_{\text{peak}} = 15 \%$$

coincidences: 550 & 915 keV

^{148}gPm decay

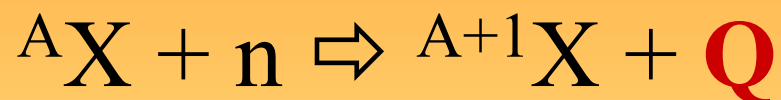


43 - 2003

Astrophysical Implications

- $N\sigma(^{148}\text{Sm}) / N\sigma(^{150}\text{Sm}) :$ **0.870 ± 0.009**
- Neutron density (this exp.): **$(4.94 \pm 0.6) 10^8 \text{ cm}^{-3}$**
- Neutron density (previous): **$(4.1 \pm 0.6) 10^8 \text{ cm}^{-3}$**

Evidence for neutron capture: **PROMPT**



$$Q = \sum \gamma_i$$

⇒ “monoenergetic” if
100 % efficiency

LANSCE @ LANL

TOF



Detector for Advanced Neutron Capture Experiments

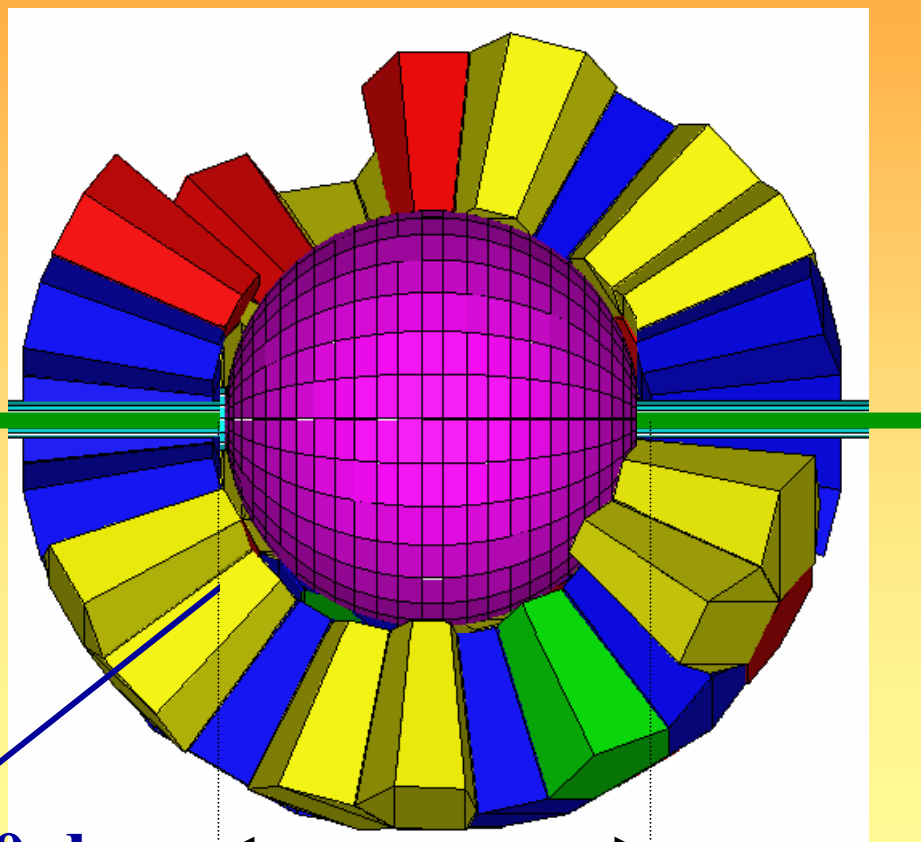
neutrons:

- spallation source
- thermal .. 500 keV
- 20 m flight path
- $3 \cdot 10^5$ n/s/cm²/decade

γ -Detector:

- 159 BaF₂ crystals
- 4 different shapes
- $R_i=17$ cm, $R_a=32$ cm
- 7 cm ⁶LiH inside
- $\epsilon_\gamma \approx 90 \%$
- $\epsilon_{\text{casc}} \approx 98 \%$

collimated
neutrons
beam



sample

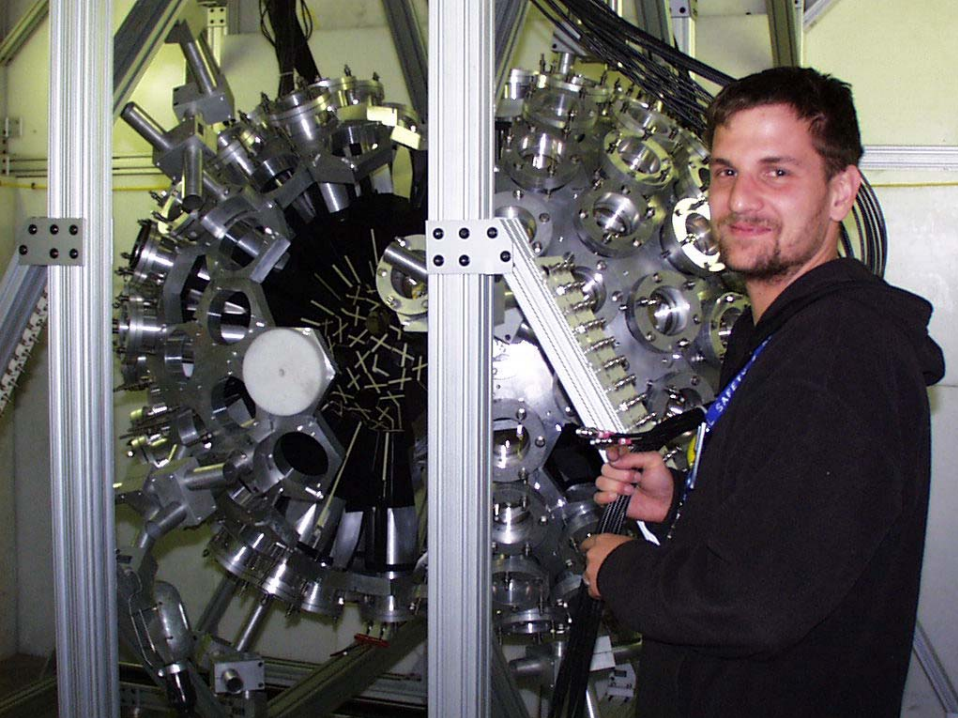
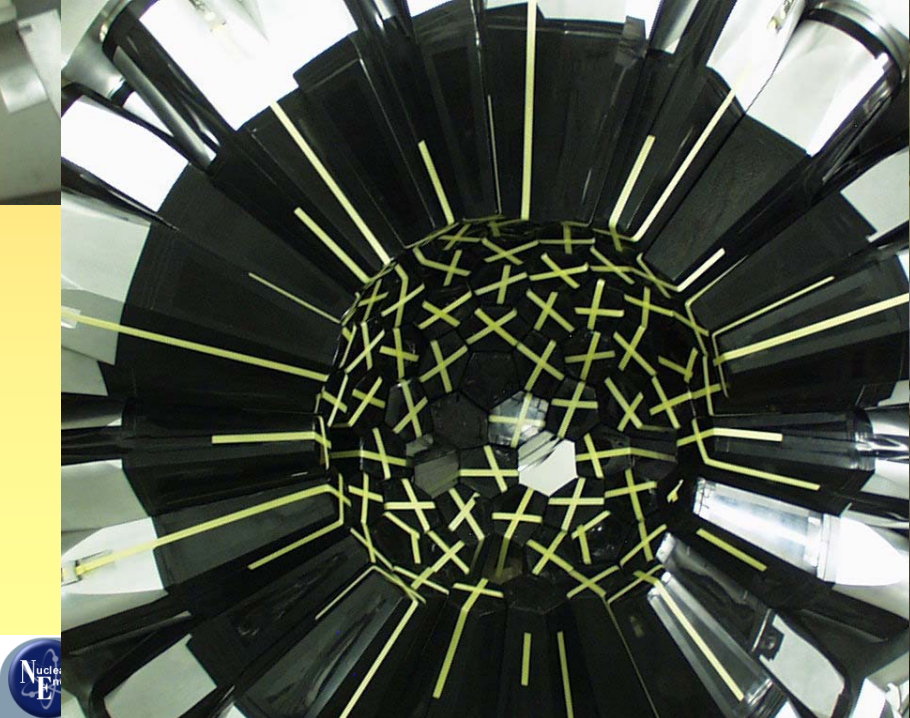
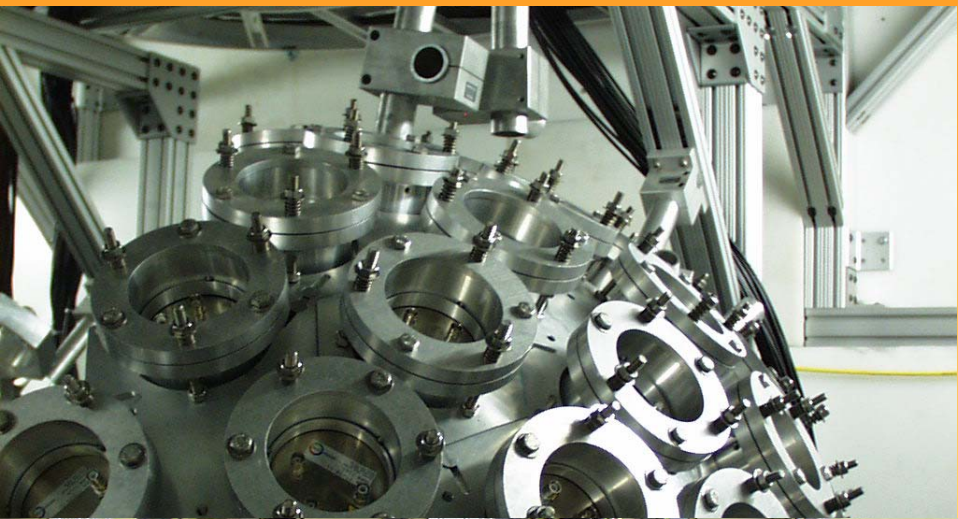
$t_{1/2} > 100$ d

$m \sim 1$ mg

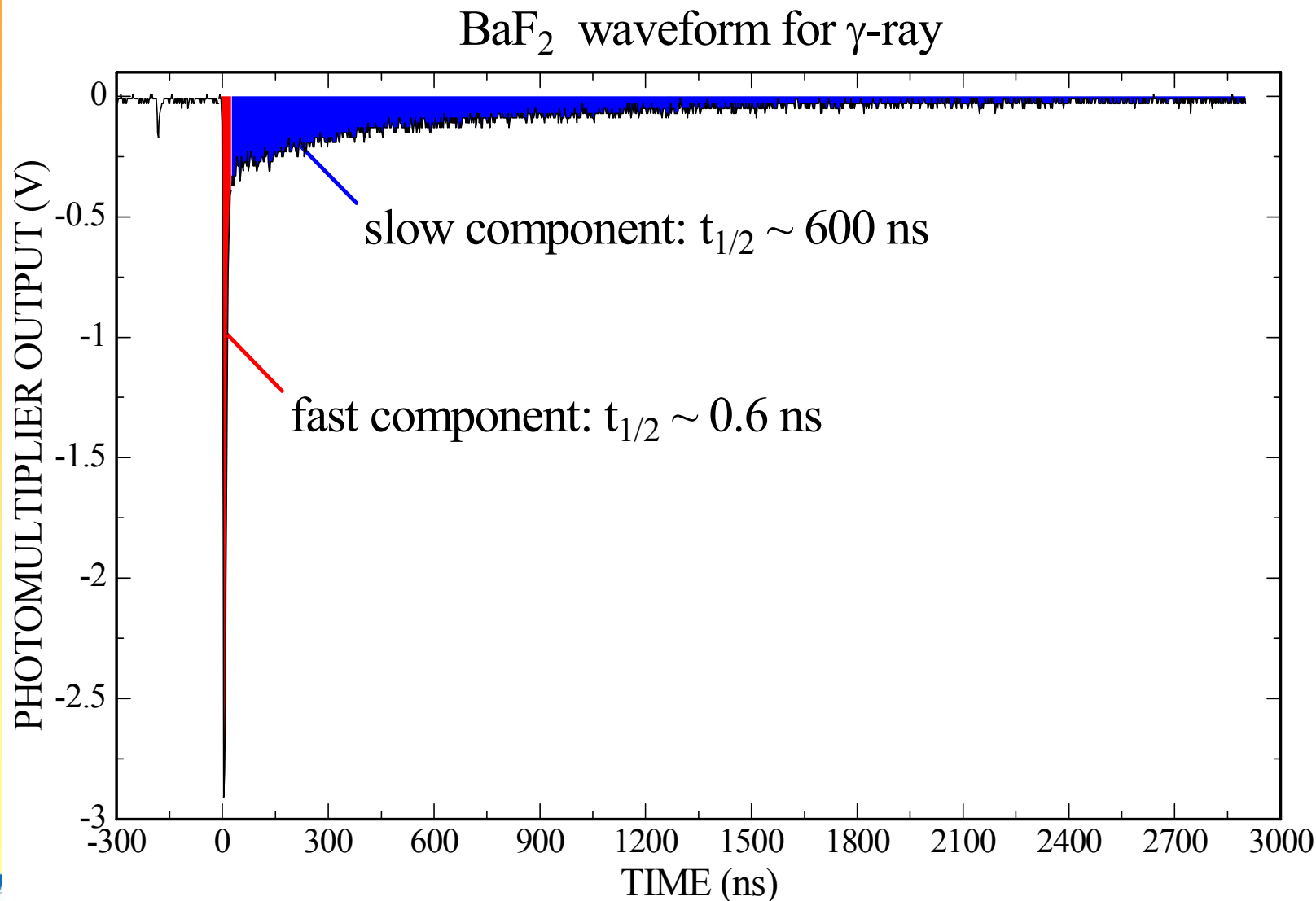
34 cm

DANCE

TOF



BaF₂ waveforms - linear



DAQ – challenges

	fast component	slow component
Timing	0.6 ns	600 ns
Energy	10 %	90 %
Amplitude	90 %	10 %

- fast component needed for timing & particle identification
- slow component needed for energy resolution

Requirements to DAQ:

- fast (~ 1 ns)
- dynamic range (~ 12 bit)
- process data within 50 ms (accelerator)
- 160 channels

DAQ – solution

12 channels high gain
12 channels low gain

Acquiris-Crate

- 14 crates
- 500 MHz PIII
- 500 MS / s
- running MIDAS frontend
- 8 bit / channel
- 2 channel / crystal
- 128000 samples memory

12 channels high gain
12 channels low gain

Acquiris-Crate

2x2 GHz PC

- running MIDAS experiment server
- ROOT interface
- 2.4 TB data storage

Total:
184 channels for fast comp.
184 channels for slow comp

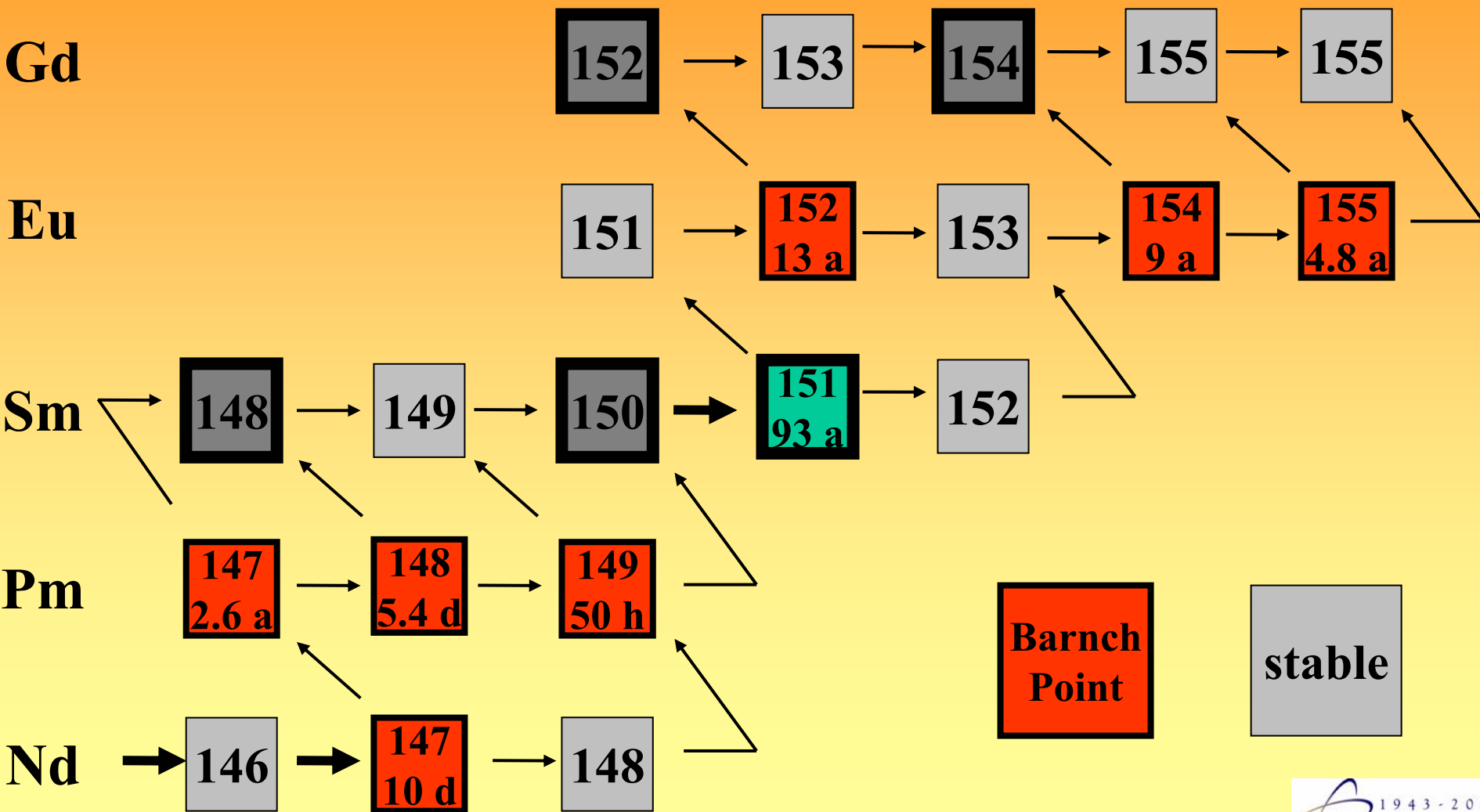
DAQ - works



First experiments with full array done!

- 320 channels FADC (160 crystals)
- 500 MS/s
- data storage on regular HD
- total file size < 2GB (OS limitation)
- event builder works!!
- problems with synchronization between crates (corrected offline)

Problem with ^{151}Sm



The early conclusion

“Obviously, none of the existing data sets does result in satisfactory solutions, a dilemma that underlines the importance of the cross sections for the unstable branch point nuclei

^{147}Pm , ^{148}Pm , and ^{151}Sm .

Therefore, vigorous efforts are necessary towards direct measurements on these radioactive isotopes.”

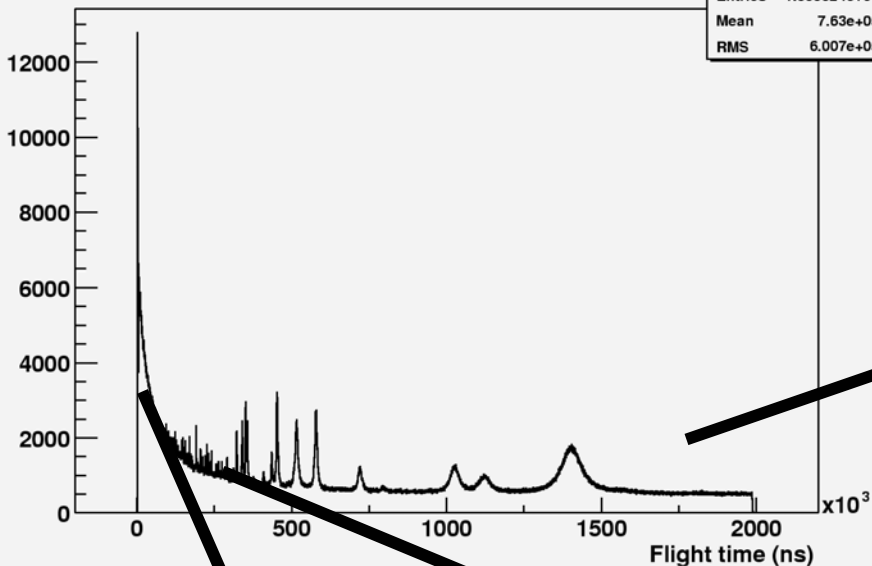
Best et. al. PRC 64, 015801, 2001

0.5 mg of $^{151}\text{Sm}(n,\gamma)$ – TOF

TOF

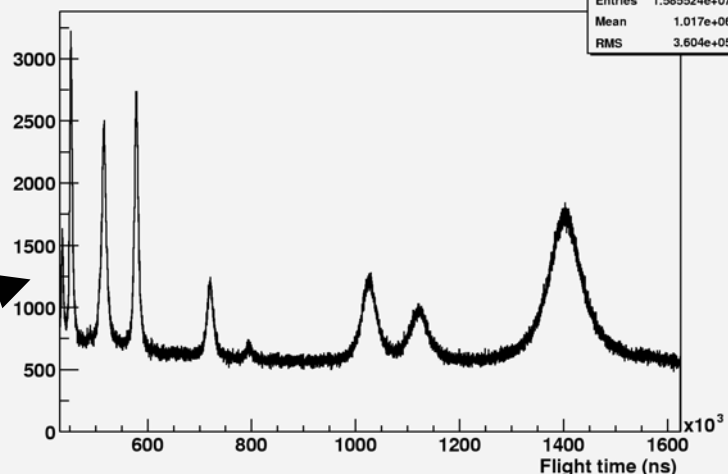
Neutron Flight Time

NTime	
Entries	1.585524e+07
Mean	7.63e+05
RMS	6.007e+05



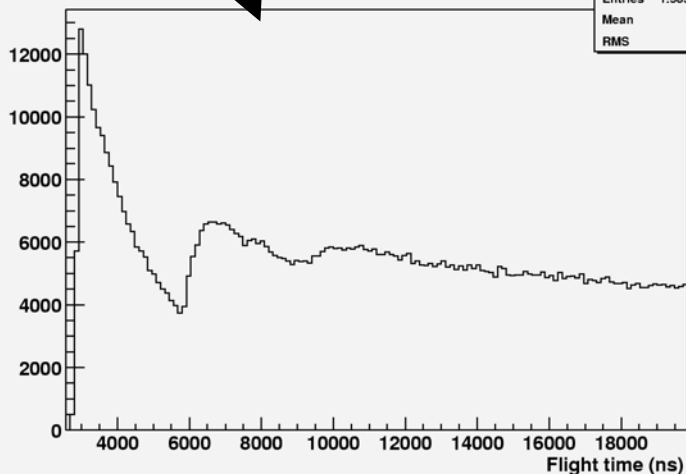
Neutron Flight Time

NTime	
Entries	1.585524e+07
Mean	1.017e+06
RMS	3.604e+05



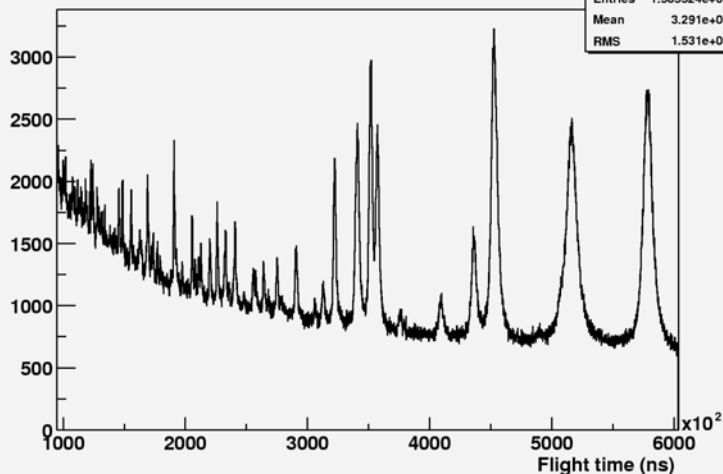
Neutron Flight Time

NTime	
Entries	1.585524e+07
Mean	1.06e+04
RMS	5007



Neutron Flight Time

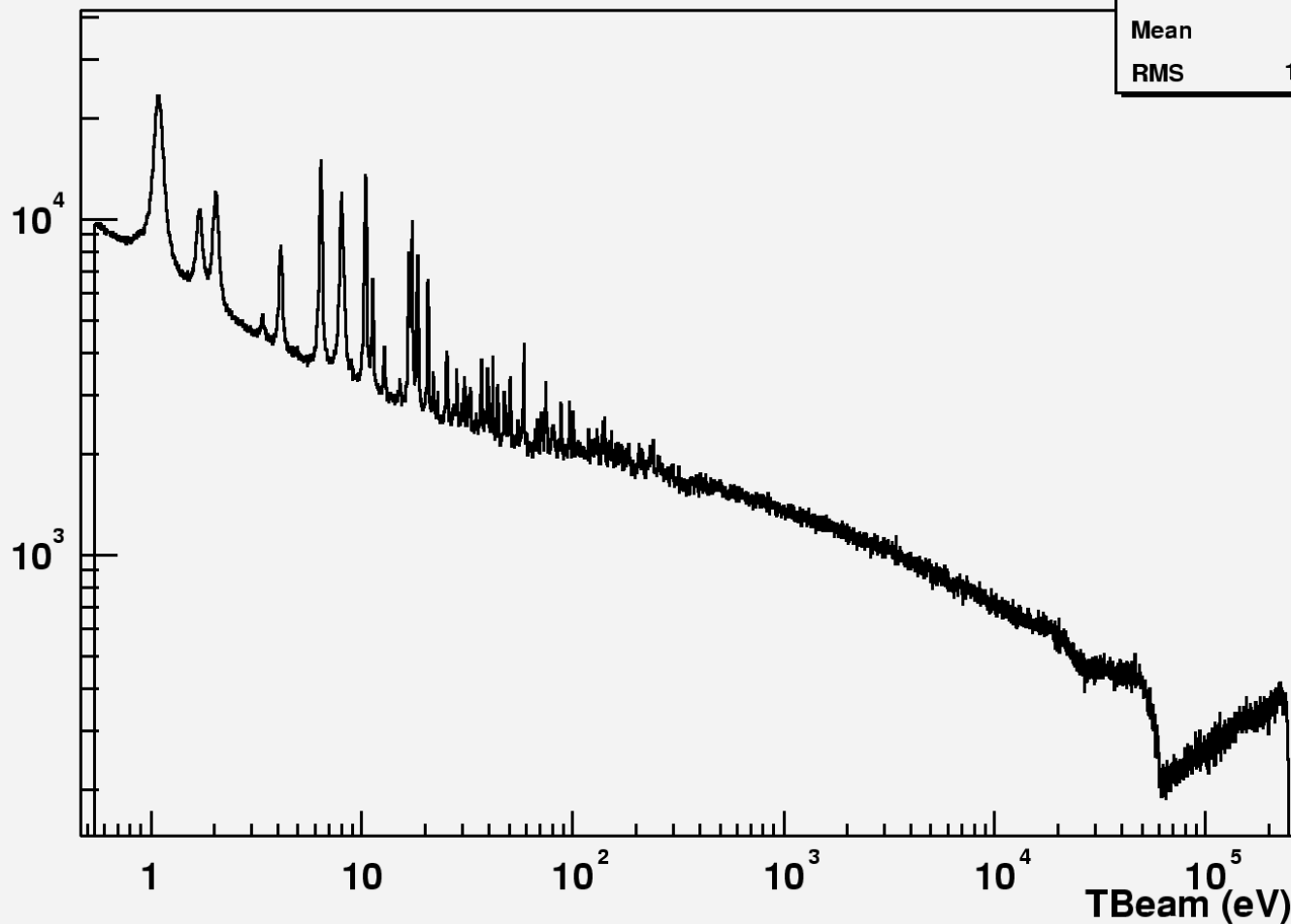
NTime	
Entries	1.585524e+07
Mean	3.291e+05
RMS	1.531e+05



$0.5 \text{ mg of } ^{151}\text{Sm}(n, \gamma) - E_n$

TOF

Beam Energy



TBeam

Entries	1.585524e+07
Mean	2603
RMS	1.694e+04

Program

Outlook

Isotope production at ILL, 2002

Target	Product	Mass	Half life
^{146}Nd	^{147}Pm	3.8 mg	2.6 yr
^{154}Sm	^{155}Eu	8.9 mg	4.8 yr
^{170}Er	^{171}Tm	9.4 mg	1.9 yr

s-process Branch Points, $t_{1/2} > 1$ yr

^{151}Sm , ^{147}Pm (FY2003)

^{63}Ni , ^{79}Se , $^{81,85}\text{Kr}$, ^{93}Zr , ^{99}Tc , ^{134}Cs , ^{135}Cs ,
 ^{152}Eu , ^{154}Eu , ^{155}Eu , ^{163}Ho , ^{171}Tm , ^{176}Lu , ^{179}Ta ,
 ^{185}W , ^{186}Re , ^{193}Pt , ^{204}Tl , ^{205}Pb

s-branch Point	before DANCE	with DANCE
^{63}Ni	-	✓
^{79}Se	-	✓
$^{81,85}\text{Kr}$	-	✓
$^{134,135}\text{Cs}$	-,~	✓
^{147}Nd	-	-
$^{147,148}\text{Pm}$	~, -	✓, -
^{151}Sm	✓	✓
^{153}Gd	-	-, ✓
$^{154,155}\text{Eu}$	-,~	✓
^{160}Tb	-	✓
^{163}Ho	-	✓
$^{170,171}\text{Tm}$	-, ✓	✓
^{179}Ta	-	✓
^{185}W	-	✓
^{204}Tl	-	✓

What else?

radioactive ion-beams – 1: surrogate reaction

$${}^AX(n,\gamma){}^{A+1}X \sim {}^AX(d,p){}^{A+1}X$$

inverse kinematics



$$d({}^AX, {}^{A+1}X)p$$

needs to be tested and very well understood

What's next?

Outlook

radioactive ion-beams - 2

RIA:

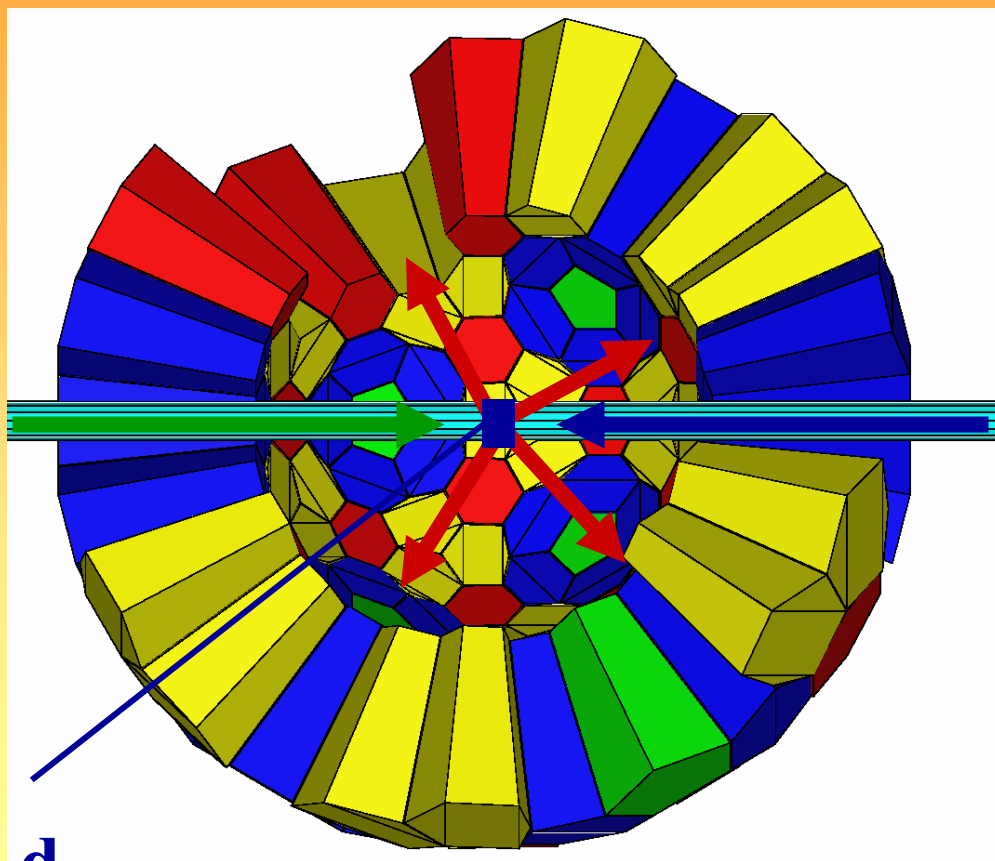
- spallation source
- $> 10^{10}$ particles / s
- 10 d

neutrons:

- VdG
- 1 .. 500 keV
- 20 d

γ -Detector:

- 4π
- $\varepsilon \approx 100\%$



sample

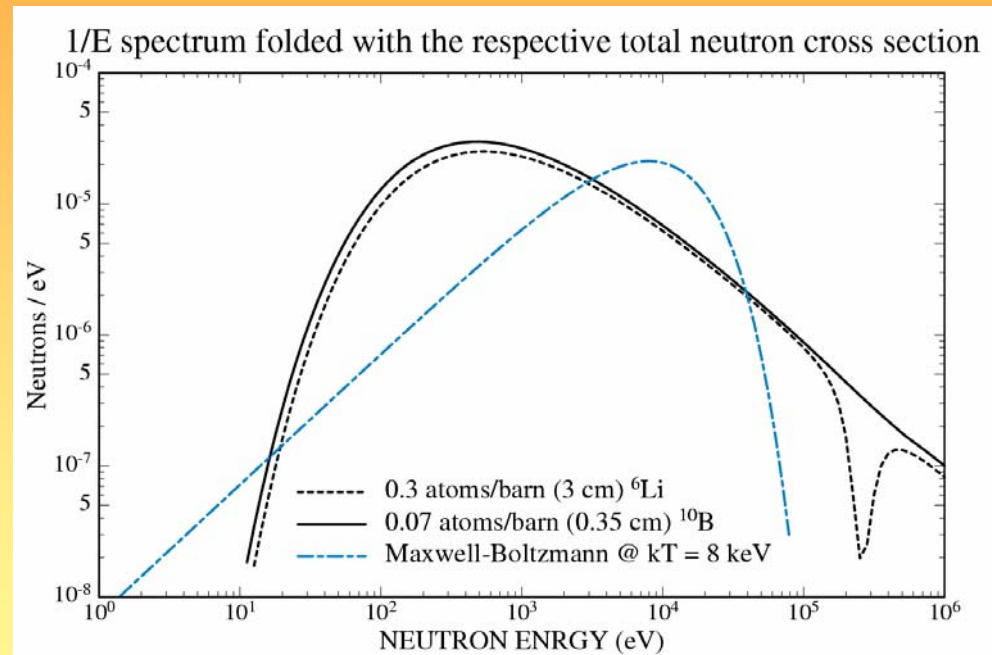
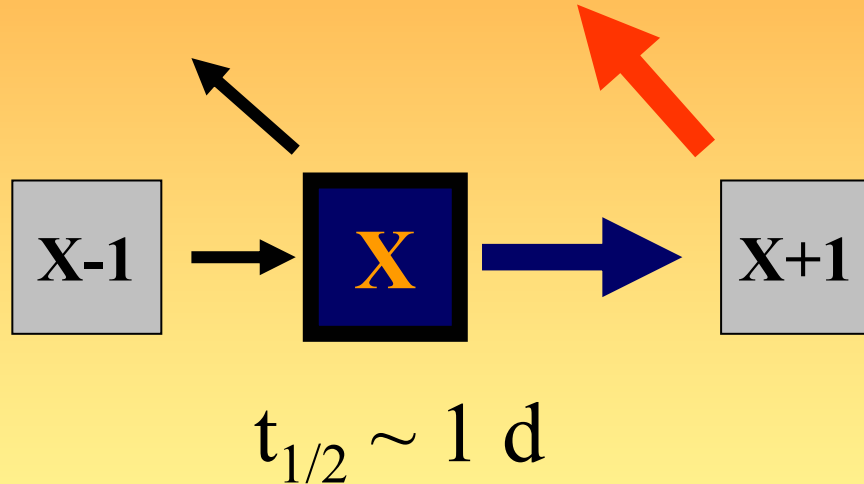
$t_{1/2} \sim 10$ d

$m \sim 10$ μ g

What's next?

Outlook

“advanced activation” close to spallation targets



Neutron spectrum needs to be “shaped”

Thanks to:

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