The Final Days of Burning Workshop at UC Santa Barbara, March 9-10 Program

9-Mar-06

Session 1	Massive Stars	BILDSTEN		
8:30 - 8:40	Lars Bildsten	UCSB	Welcome	
8:40 - 9:10 9:10 – 9:20	Dave Arnett Discussion	University of Arizona	Primer on Massive Stars	
9:20 - 9:50	Allessandro Chieffi	Frascati, Italy	The advanced evolutionary phases of massive	
9:50 – 10:00	Discussion		stars and their influence on the final abundance yields	
10:00 - 10:20 10:20 – 10:30	Falk Herwig Discussion	LANL	Progenitor evolution of ONeMg core-collapse supernova	
10:30 - 11:00	Coffee Break			
11:00 - 11:20 11:20 – 11:30 Session 2	Marco Pignatari Discussion Weak Interaction	U. Torino, Italy AUSTIN	The Weak s-process after Core he Burning: the convective shell C-burning contribution	
11:30 – 12:00 12:00 – 12:10	Karlheinz Langanke Discussion	TU-Darmstadt, Germany	Primer on Weak Interaction Physics	
12:10 - 1:30	Lunch in Kohn Hall Court Yard			
1:30 – 1:50 1:50 – 2:00	Gabriel Martinez- Pinedo Discussion	GSI Darmstadt, Germany	Electron Capture on Core Collapse Supernova	
2:00 - 2:20 2:20 - 2:30	Raph Hix Discussion	ORNL	The Role of Nuclear Electron Capture during Stellar Core Collapse	
2:30 – 2:50 2:50 – 3:00	Remko Zegers Discussion	NSCL/MSU	Charge-exchange reactions as a tool to extract weak rates for astrophysics	
3:00 - 3:20	Coffee Break			

Session 3 Nuclear Reaction Processes (SCHATZ)

3:20 - 3:50	Michael Wiescher	University of Notre	Nuclear Reactions in Stellar Burning density stellar environments
3:50 - 4:00	Discussion	Dame	
4:00 - 4:20 4:20 - 4:30	Christian Diget Discussion	U. Arhus, Danmark	Status of the Triple Alpha reaction
4:30 - 4:50	Carl Brune	Ohio University	The Stellar 12C+alpha Fusion Rate: Present
4:50 - 5:00	Discussion		Uncertainties and Prospects for their Reduction
5:00 - 5:30	Collective Discussion (led by Bildsten and Schatz)		

6:30 Dinner at El Paseo in Downtown Santa Barbara

10-Mar-06

The second day will be devoted to discussions on topics relevant for the understanding and interpretation of late stellar pre-supernova evolution. The following topics have been suggested, but more topics can be identified during the collective discussion after the third session of the first day.

8:30 am

1.	Triple α and ${}^{12}C(\alpha,\gamma)$ rates Do we have to know more than the properties of the Hoyle State? For SNII, is absolute value of the 3 alpha rate important or only its ra to ${}^{12}C(\alpha,\gamma)$?	Chair: Sam Austin tio			
2.	What is the origin of ⁶⁰ Fe? What is the importance as a diagnostics of a supernova explosion	Chair: Alessandro Chieffi			
3.	Is there a possible set(small) of properties that characterize a progenitor as regards to the ensuing explosion	Chair: Dave Arnett			
10:00-10:30: COFFEE					
4.	What are the best cases for validating model calculations for B(GT) in the A>60 mass range; what measurements are possible	Chair: Karlheinz Langanke			
5.	What are the best cases for validating model calculations of forbidden strength; what measurements are possible.	Chair: Remco Zegers			
6.	How reliable are calculations of neutrino cross sections?	Chair: Gail Mclaughlin			
12:00 Lunch					